

# ORIGINS

Space Telescope

UCI

**Asantha Cooray**

**@acooray**

<http://origins.ipac.caltech.edu>

Twitter: @NASAOriginsTele





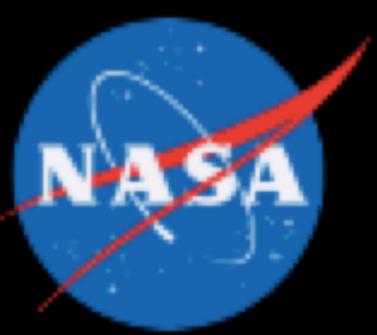
## Study Team

- **Study Chairs:** Asantha Cooray, *UC Irvine*; Margaret Meixner, *STSCI/JHU*
- **Study Scientist:** David Leisawitz, *GSFC*
- **Deputy Study Scientist:** Johannes Staguhn, *GSFC/JHU*
- **Study Manager:** Ruth Carter, *GSFC*
- **NASA HQ Program Scientists:** Kartik Sheth, Dominic Benford

- **NASA Appointed Members:** Lee Armus, *IPAC*; Cara Battersby, *CfA*; Edwin Bergin, *Michigan*; Matt Bradford, *JPL*; Kim Ennico-Smith, *Ames*; Lisa Kaltenegger, *Cornell*; Gary Melnick, *CfA*; Stefanie Milam, *GSFC*; Desika Narayanan, *University of Florida*; Klaus Pontopiddan, *STSCI*; Alexandra Pope, *UMass*; Thomas Roellig, *Ames*; Karin Sandstrom, *UCSD*; Kevin Stevenson, *STScI*; Kate Y. L. Su, *Arizona*; Joaquin Vieira, *UIUC*; Edward Wright, *UCLA*; Jonas Zmuidzinas, *Caltech*
- **Ex-officio representatives:** Susan Neff & Deborah Padgett, *NASA Cosmic Origins Program Office*; Susanne Alato, *SNSB*; Douglas Scott, *CAS*; Maryvonne Gerin, *CNES*; Itsuki Sakon, *JAXA*; Frank Helmich, *SRON*; Roland Vavrek, *ESA*; Karl Menten, *DLR*; Yong-Seon Song, *KASI*; Sean Carey, *IPAC*
- **NASA Study Center (Goddard Space Flight Center) Team:** Anel Flores (Mission Systems Engr), James Kellogg (Instrument Systems Engr), Michael DiPirro (Chief Technologist), Louis Fantano (Thermal Systems Engr), Andrew Jones (Mechanical Systems Engr), Joseph Howard (Optical Systems Engr), James Corsetti (Optical Engr), Ed Canavan (Cryo Engr), Johannes Staguhn (Instrument Scientist)
- **Study Advisory Board:** Jon Arenberg, *Northrup Grumman*; John Carlstrom, *Chicago*; Harry Ferguson, *STScI*; Tom Greene, *Ames*; George Helou, *IPAC*; Charles Lawrence, *JPL*; Sarah Lipsky, *Ball Aerospace*; John Mather, *GSFC*; Harvey Moseley, *GSFC*; George Rieke, *Arizona*; Marcia Rieke, *Arizona*; Jean Turner, *UCLA*; Meg Urry, *Yale*.



Tracing the rise of dust & metals in galaxies  
and the path of water across cosmic time to  
Earth and other habitable planets.



## How did we decide on the Origins Space Telescope?

(1) Define Science  
*(consider 2018-2035  
science developments;  
science goals for 2035)*

(2) Prioritize science  
*(STDT internal voting  
process - completed  
August 2016 meeting)*

(3) Derive mission and  
instrument requirements  
*(Completed Nov 2016  
meeting)*

Science process is through **six science working groups (SWGs)**. **Membership in SWGs is open to the community.**

OST interest groups in Europe and Japan

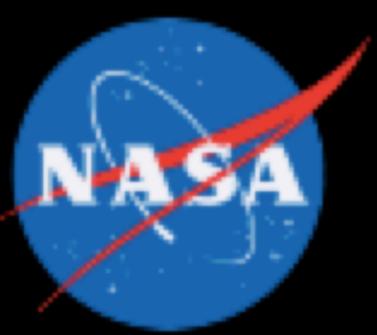
*(about 250 community members active already! SWG listings on our website)*



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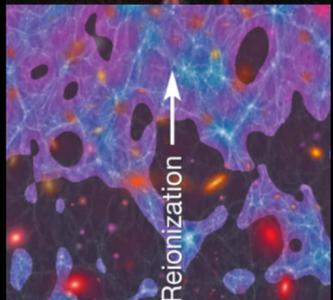
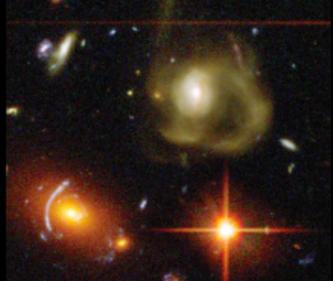
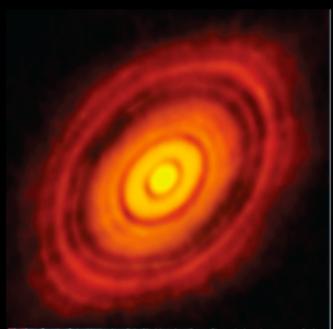
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## Science Working Groups

- **Solar System:** Stefanie Milam
- **Exoplanets:** Lisa Kaltenegger and Kevin Stevenson
- **Disks and Protoplanets:** Klaus Pontoppidan and Kate Su
- **Milky Way, ISM and Nearby Galaxies:** Karin Sandstrom and Cara Battersby
- **Galaxy Evolution over Cosmic Time:** Lee Armus and Alex Pope
- **Early Universe and Cosmology:** Matt Bradford and Joaquin Vieira



**Science Case - Number and Title**

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19: The Rise of Metals

14: Biosignatures of Transiting Exoplanets \*\*revamped\*\*

27: The First Dust

9: Water Content of Planet-Forming Disks

21: Connection Between Black Hole Growth and Star Formation Over Cosmic Time

15: Direct Detection of Protoplanetary Disk Masses

26: Birth of Galaxies During Cosmic Dark Ages

18: Galaxy Feedback from SNe and AGN to  $z \sim 3$

30: Survey of Small Bodies in the Outer Solar System

16: Direct Imaging of Exoplanets \*\*revamped\*\*

22: Star Formation and Multiphase ISM at Peak of Cosmic Star Formation

29: Thermo-Chemical History of Comets and Water Delivery to Earth

5: Galaxy Feedback Mechanisms at  $z < 1$

4: Water Transport to Terrestrial Planetary Zone

13: Frequency of Kuiper Belt Analogues

24: Feedback on All Scales in the Cosmic Web

7: Magnetic Fields and Turbulence - Role in Star Formation

New#1 The EBL (extra-galactic background light) with OST

New#2: Determining the cosmic-ray flux in the Milky Way and nearby galaxies

17: Episodic Accretion in Protostellar Envelopes and Circumstellar Disks

8: Formation and History of Low-Mass Ice Giant Planets

32: Find Planet IX

New#3: Fundamental of dust formation around evolved stars

New#4: The dynamic interstellar medium as a tracer of galactic evolution

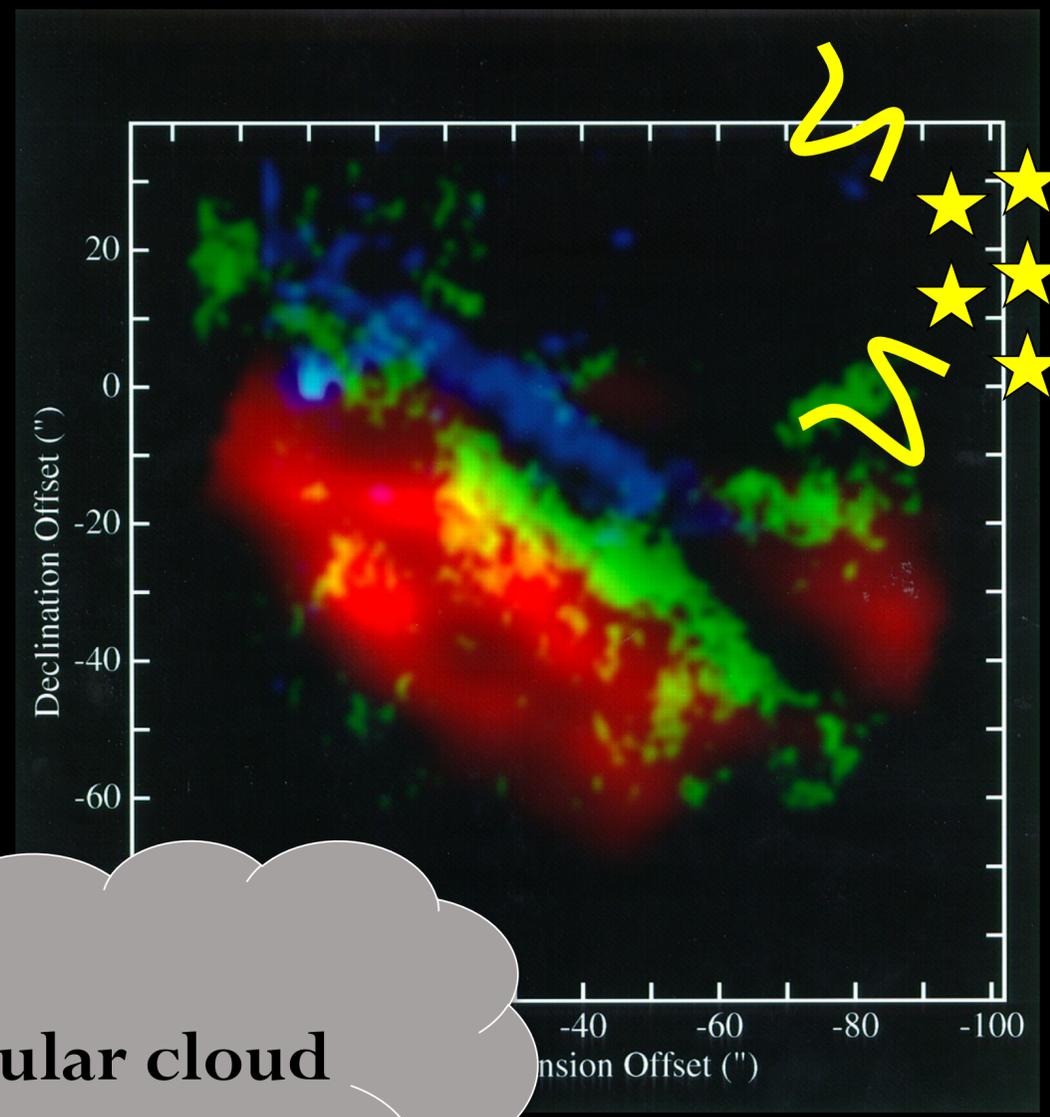
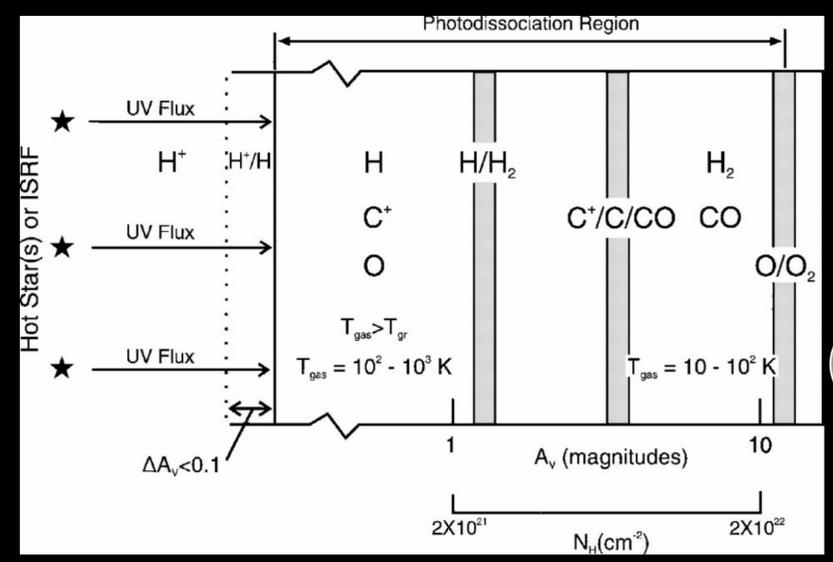
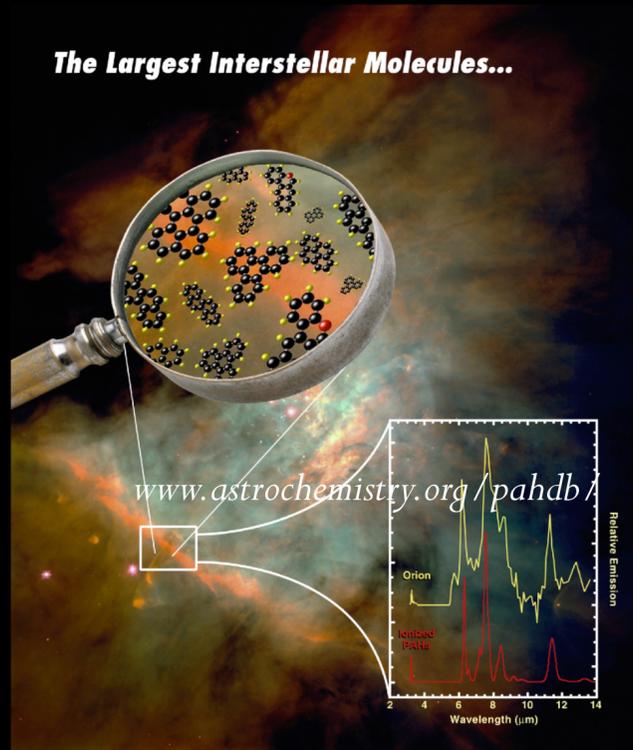
New#5: Probing magnetic fields with fine structure lines

Science white papers are publicly available from <https://asd.gsfc.nasa.gov/firs/>  
STDT will revise these and also collect more papers from community in Summer 2017

## An Example Science Traceability Matrix

OST Science Case Number/Title Theme	OST Science Theme NASA Science Goal Decadal Science Goal	Science Objectives	Science Requirements		Instrument Requirements		
			Science Observable	Measurement Requirement	Technical Parameter	Technical Requirement	Instrument(s)
<p><b>19, Rise of Metals, Dust, and the First Galaxies</b></p> <p>Trace the dust and metal enrichment history of the early Universe. Find the first cosmic sources of dust, and search for evidence of the very earliest stellar populations forming in pristine environments.</p>	<p><b>OST-2:</b> <i>(Charting the) Rise of Metals, Dust, and the First Galaxies</i></p> <p><b>NASA-2:</b> <i>How did we get here?</i></p> <p><b>Decadal-1:</b> <i>Cosmic Dawn</i></p>	<p>Trace the rise of metals and (a) determine the evolution in metallicity from <math>z=1</math> to <math>z=3</math> to 0.1 dex down to <math>10^{11}L_{\text{sun}}</math>; (b) determine the cosmic metal abundance <math>\Omega_{\text{metals}}</math> from <math>z=0</math> to <math>z=8</math> to 0.1 dex accuracy in 8 redshift bins; and (c) measure the multiple phases of the ISM to infer the physical phenomena that regulate SF efficiency at the peak of cosmic star formation at <math>z=1-3</math>.</p>	$z=1-3$ relative metallicity tracer: [NeII]12.8, [NeIII]15.6, [SIII]18.7, [SIV]10.5; $z=0-8$ relative metallicity tracer: [OIII] 52+88 $\mu\text{m}$ , [NIII] 57 $\mu\text{m}$ ; cooling and heating of the ISM through [OI], [OIII], [NII], [CII].	Rest-frame mid and far-IR spectral mapping to select $z=0$ to 8 galaxies	Wavelength range	20-800 $\mu\text{m}$	incoherent spectrometer, low res mode
			Identify galaxies in a tiered spectral mapping survey	Spatial resolution	5 arcsec at 200 $\mu\text{m}$ (min. 9 m Telescope)		
			Measure line flux densities of identified galaxies	Spectral line sensitivity	1 e-21 W m-2 (driven by the MIR lines)		
				Spectral Resolving power	$\lambda/\Delta\lambda = 500$		
				survey area, instantaneous FOV, FoR	10 deg <sup>2</sup>		
	Mapping Speed						

## How do we probe the interstellar medium in high redshift galaxies?



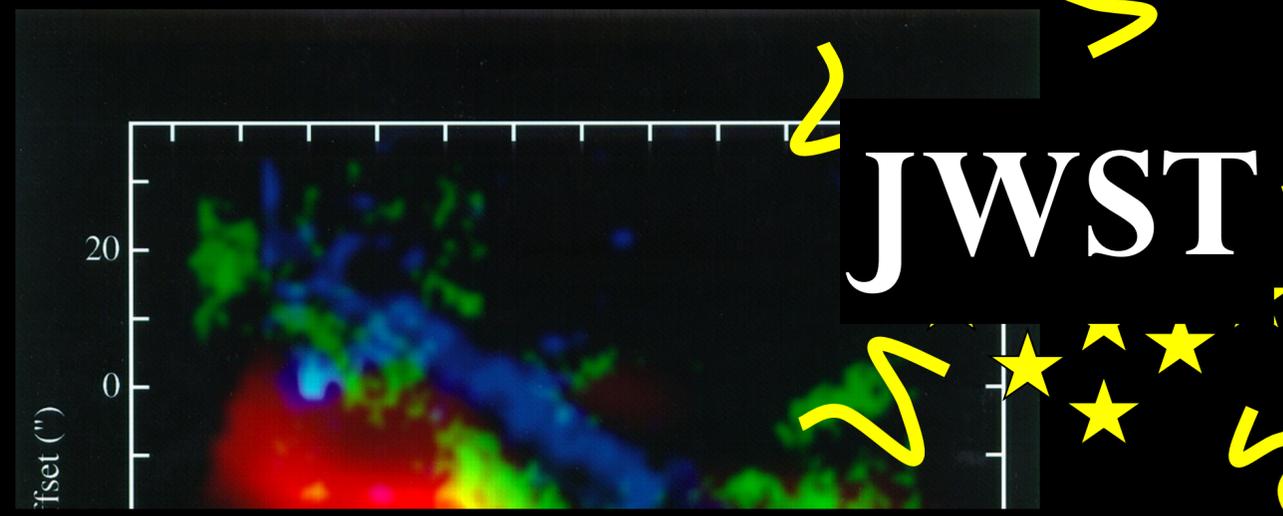
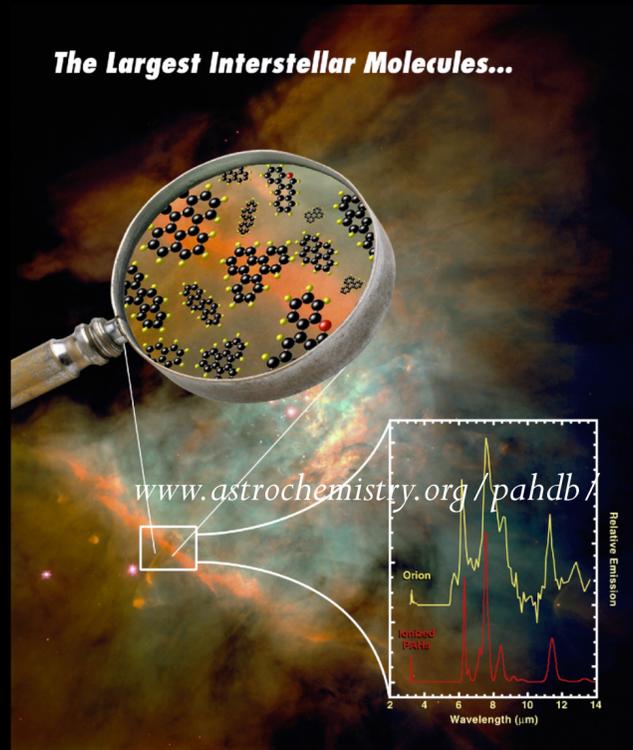
PAH

$H_2$

CO

Molecular cloud

## How do we probe the interstellar medium in high redshift galaxies?



# Origins Space Telescope



H<sub>2</sub>  
CO



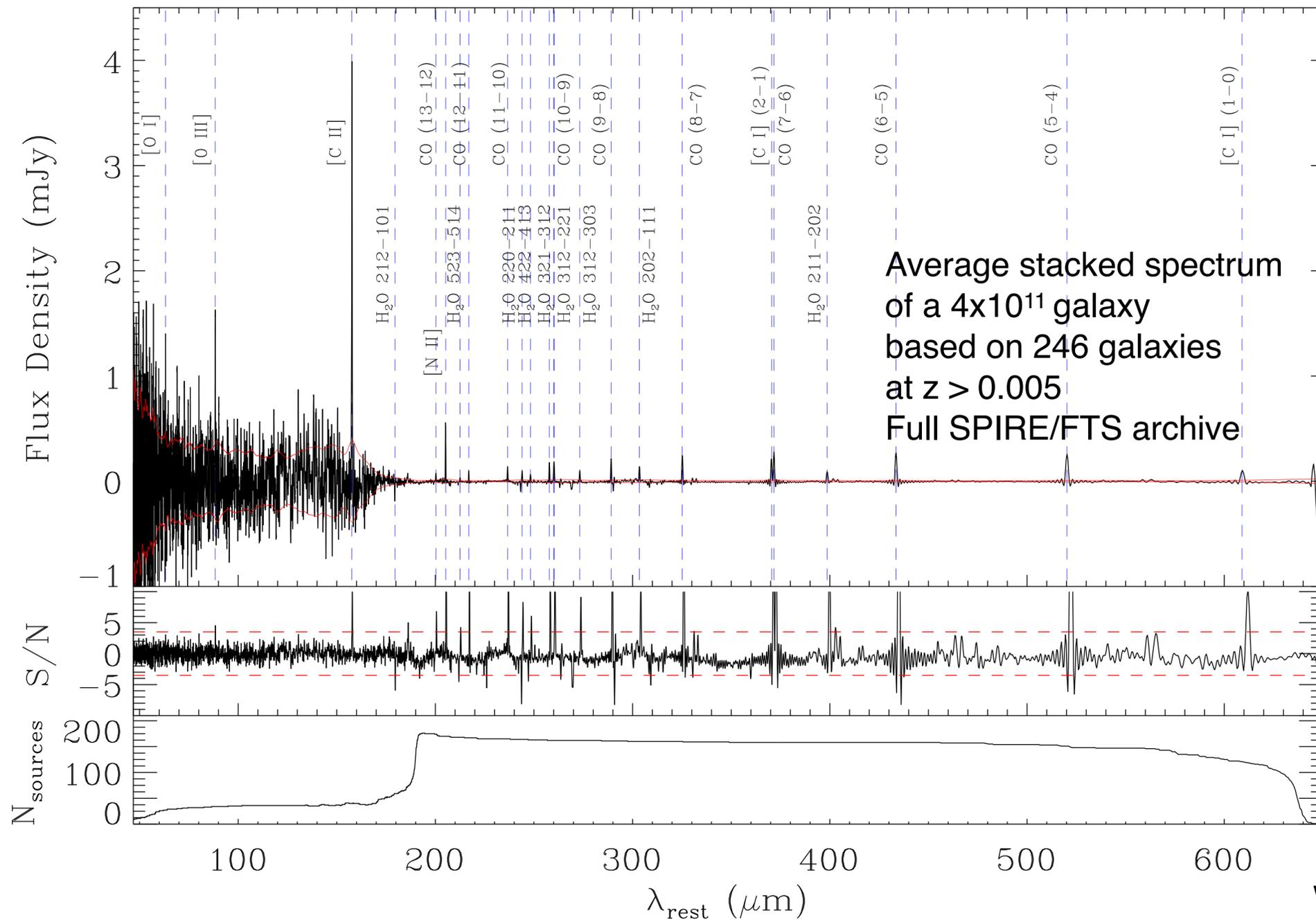
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## Infrared is rich in key spectral lines!



Wilson et al. 2017



ALMA



JWST



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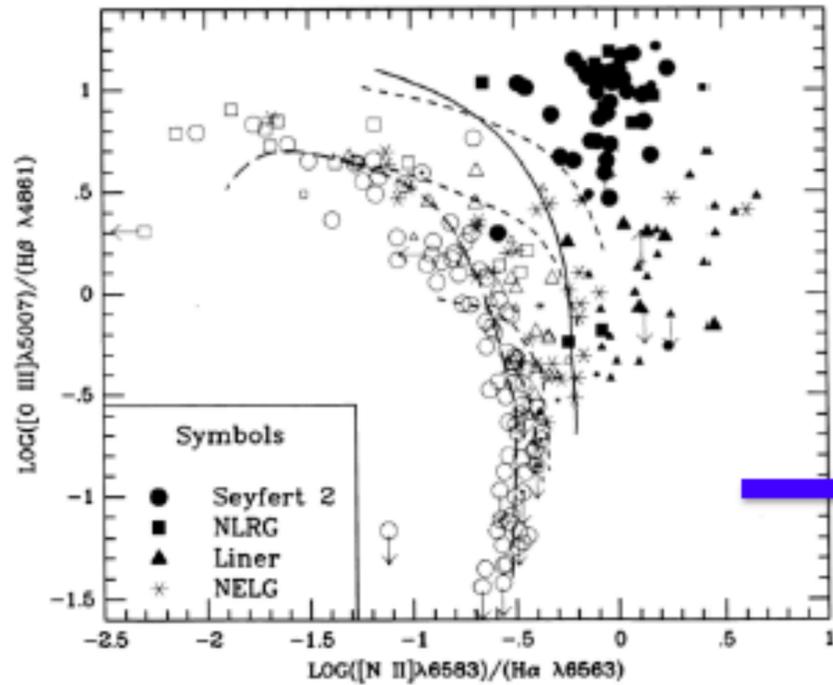
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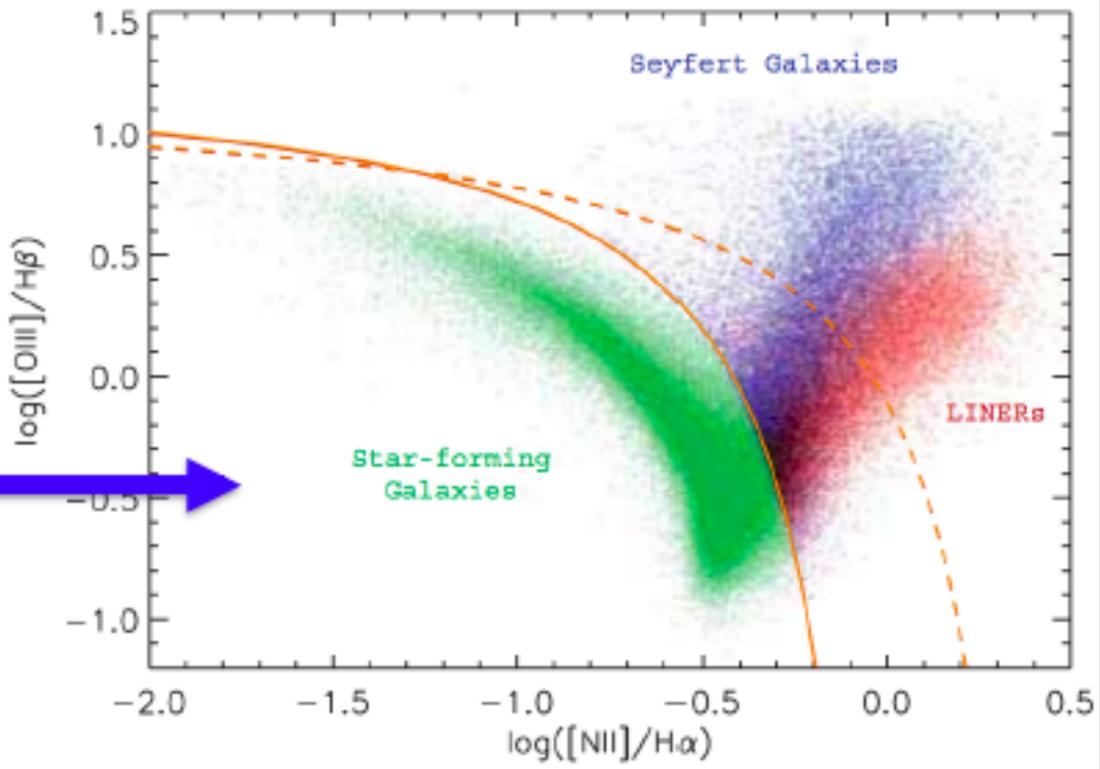


OPTICAL SPECTROSCOPY

Veilleux & Osterbrock 1987 (~100 galaxies)

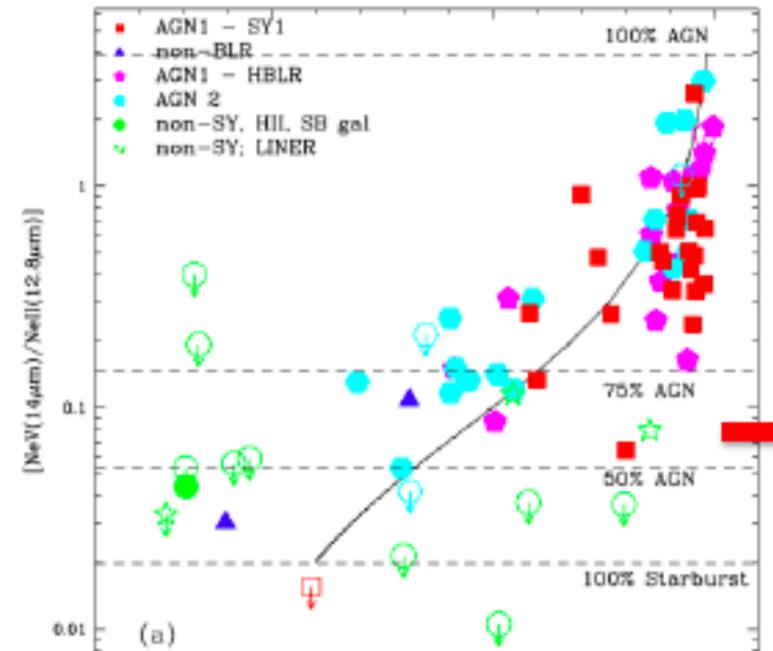


Groves+ 2006 (>10^5 galaxies)



MIR-FIR SPECTROSCOPY

Tommasin+ 2010 (~60 galaxies)



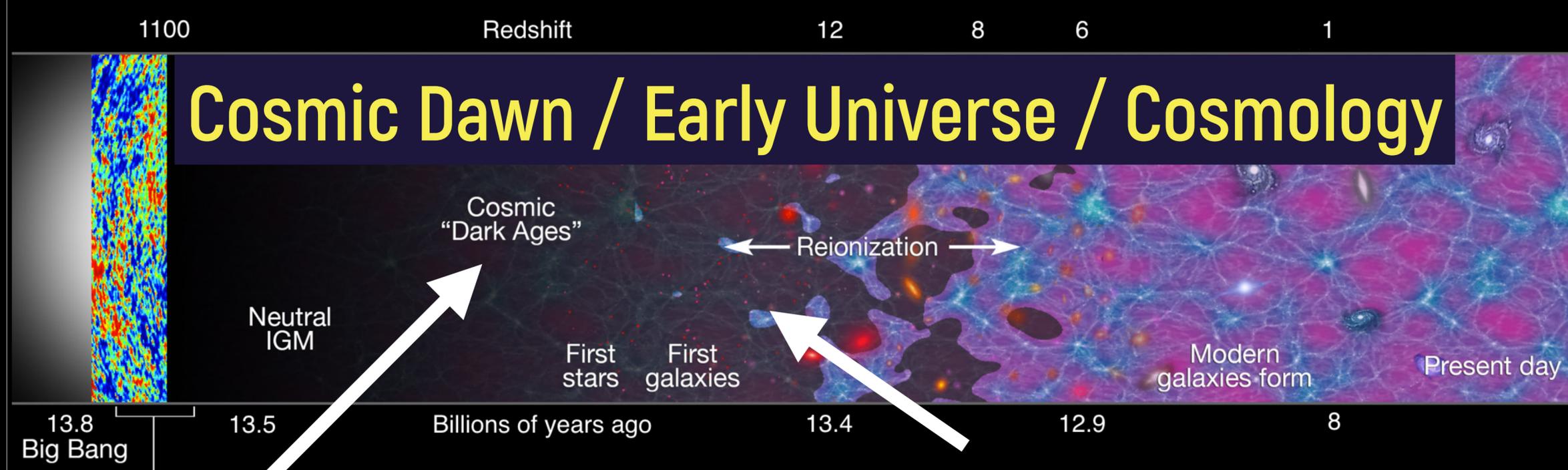
**Origins Space Telescope:**  
~ mid and far-IR spectra for 10<sup>6</sup> galaxies, from starbursts to Milky Way-like galaxies (2030+ -> 20 year development consistent with optical technology development to get million optical spectra)





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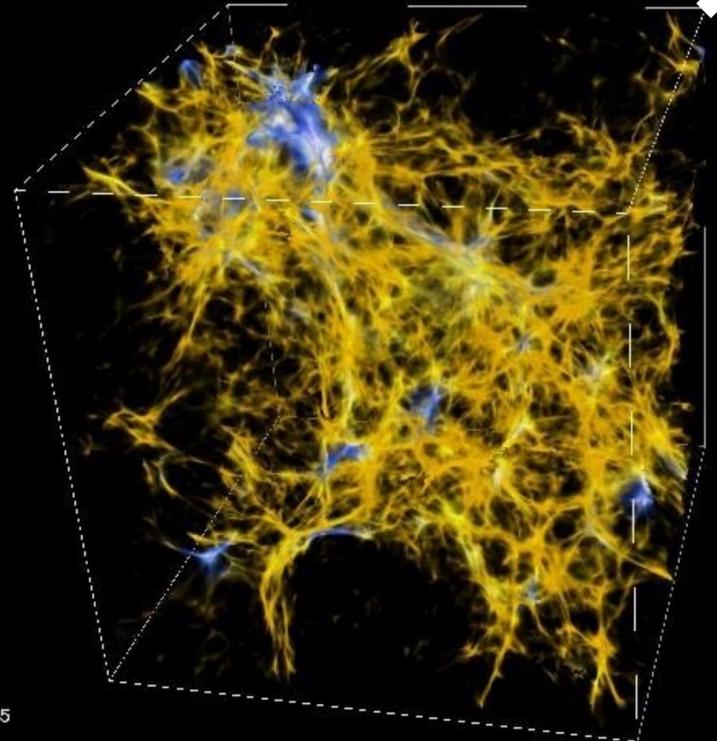
## Cosmic Dawn / Early Universe / Cosmology

Origins goes further!

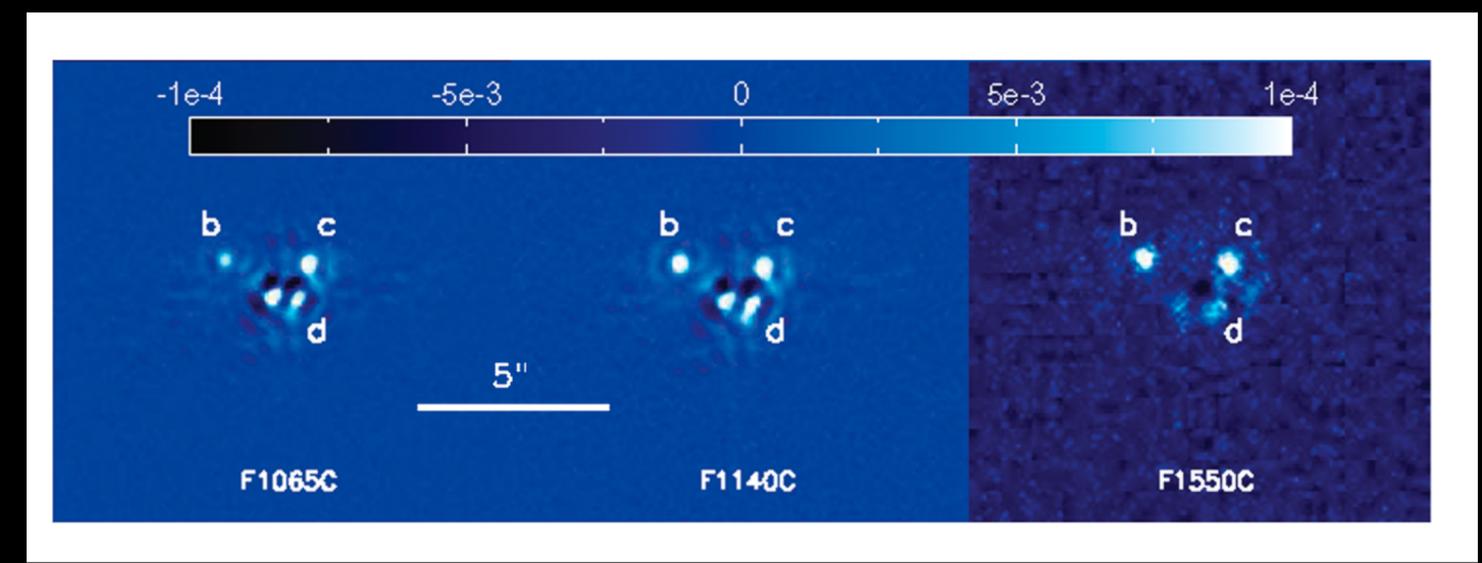
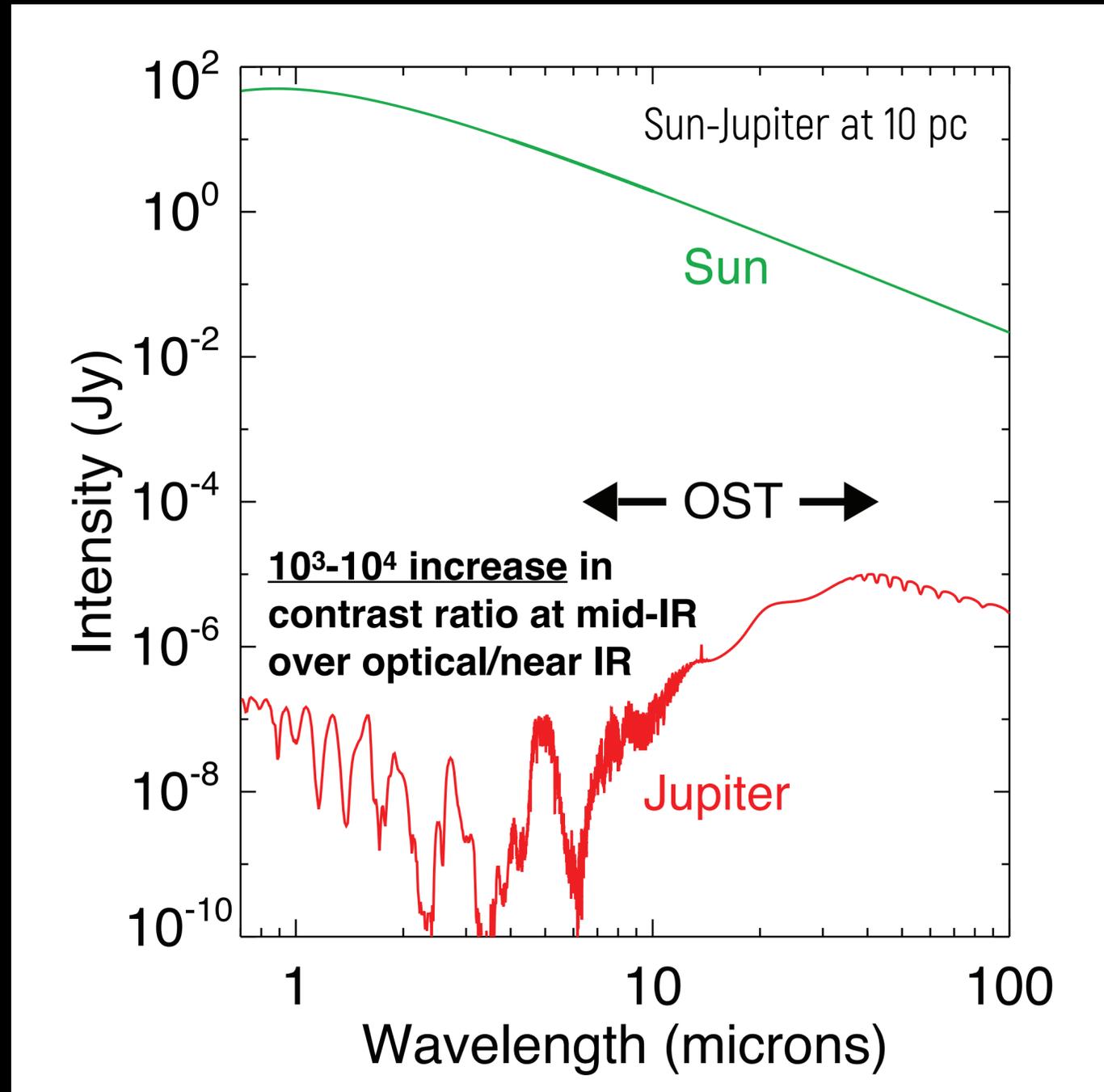
JWST/WFIRST capability is detecting first stellar emission

- ◆ Origins Space Telescope will venture beyond JWST and image gas collapsing to form first stars!
  - Primordial cooling via H<sub>2</sub> rotational lines
  - Seeds of super massive black holes

To detect primordial H<sub>2</sub> line cooling at formation sites of first stars and galaxies at  $z \sim 10-15$  *Origins Space Telescope* sensitivity will need to be down to  $10^{-23}$  Wm<sup>-2</sup> in a deep field integration in rotational lines (rest-frame 12.3, 17, 28  $\mu$ m)



## Weather and climate on exoplanets: Super Earths to Jupiters

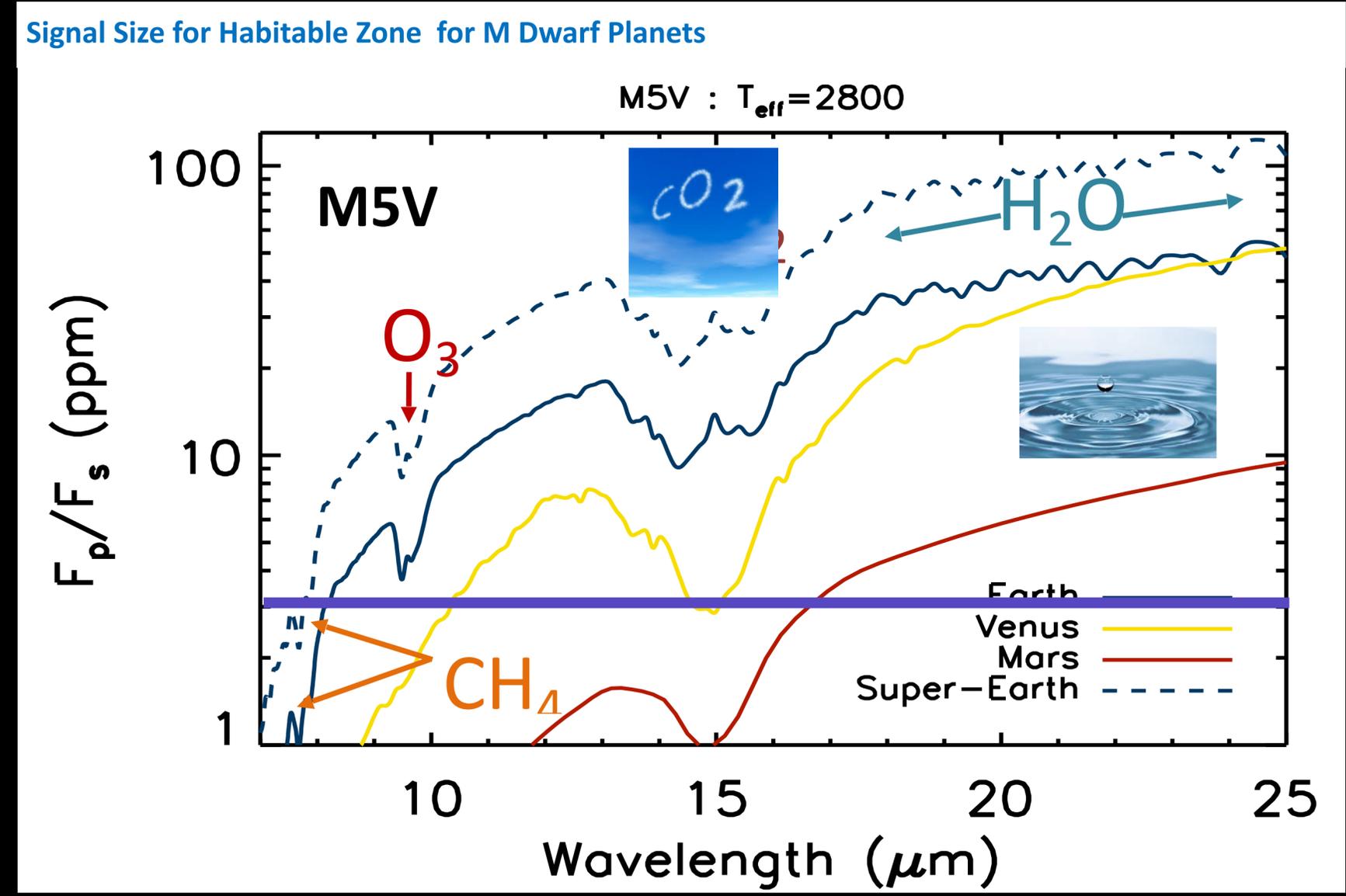
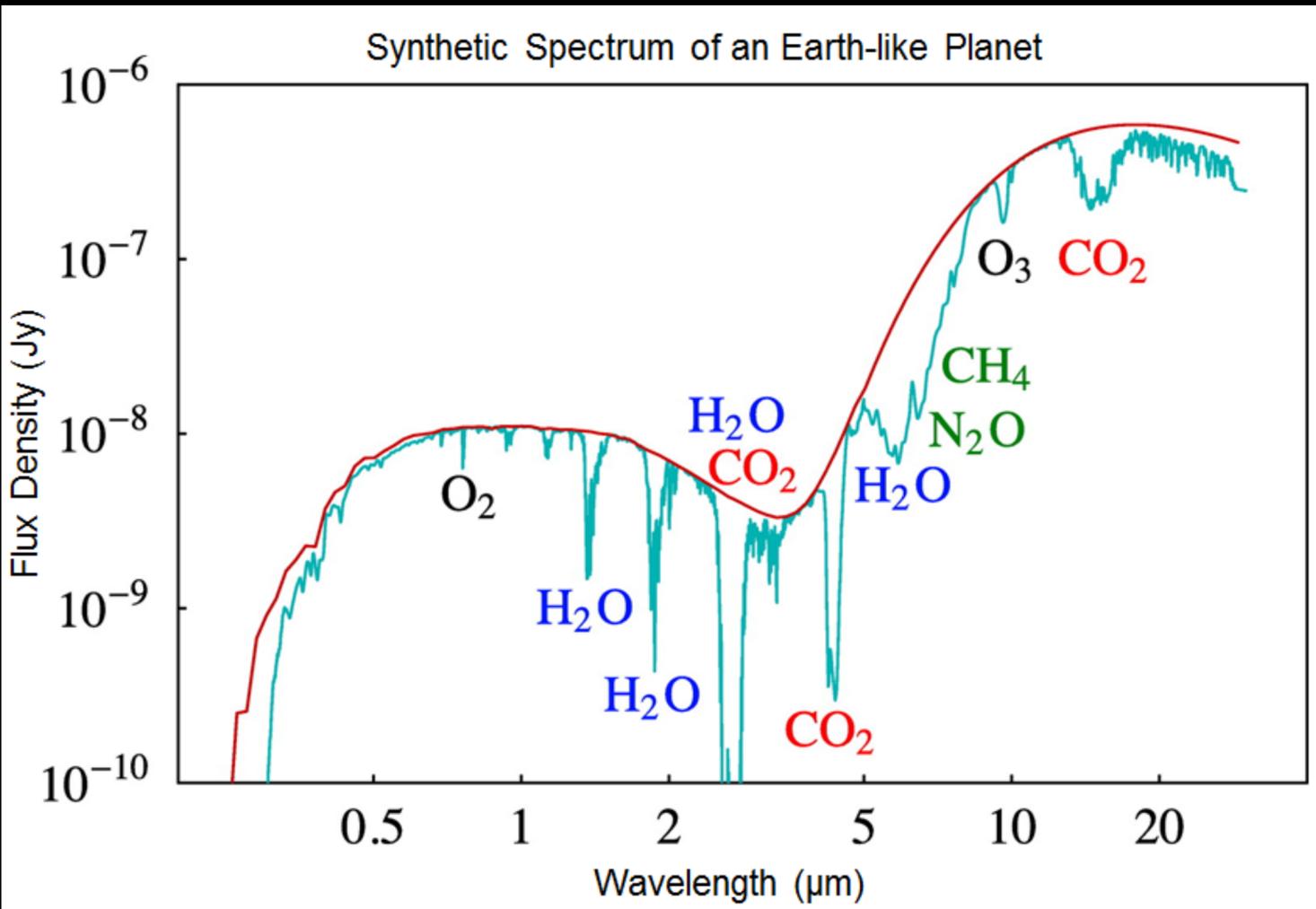


Simulated MIRI coronagraphy of HR 8799

**Transit/secondary eclipse spectroscopy (Earths/super-Earths around MK-dwarfs)**

**Direct imaging via a mid-IR coronagraph (Jupiters/Saturns/Neptunes)**

## Weather and climate on exoplanets: Super Earths to Jupiters

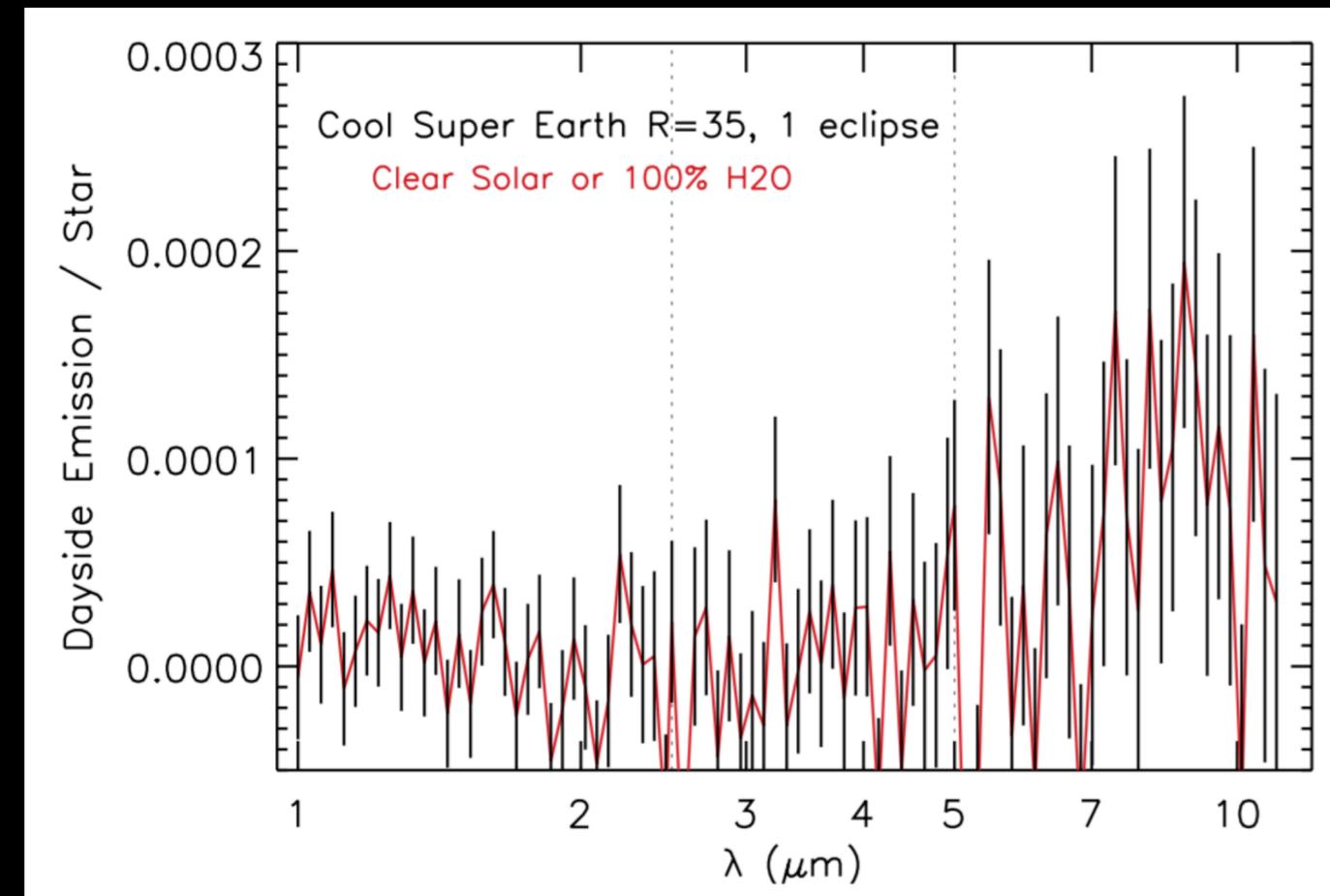


Rauer et al. 2011 Potential Bio-Signatures in Super-Earth Atmospheres, A&A

## To detect biosignatures:

- Spectral resolving power ( $\lambda/\Delta\lambda$ ) of 30-50
- Noise floors < 10 ppm
  - (M3V@20 pc – 2 hr at 7  $\mu\text{m}$ )
- Key spectral signatures of Super-Earths that Origins will detect:
  - 9  $\mu\text{m}$  for ozone (biosignature)
  - 7  $\mu\text{m}$  for methane (life detection)

*Origins Space Telescope will have mid-IR capability down to 6  $\mu\text{m}$ ; noise floor will be due to mid-IR detector stability.*



**At 50ppm JWST cannot study habitable zone worlds (Greene et al. 2016)**

## Mission Study Design Implementation

- Led at GSFC, coordinated through STDT via the Mission Development Working Group (led by Tom Roellig).
- We have formed five instrument teams (ITs), two foreign contributions (Europe through CNES and Japan/JAXA)
- Formal industry partnership through a **Cooperative Agreement Notice (CAN)** with **Ball Aerospace as the lead, Northrop Grumman, Harris, and Lockheed Martin** as contributors. Industry is studying items put forward by the STDT and Study Center (thermal, deployment, Sun shield/Solar panels).



## Origins Space Telescope Study Plan

- June 2017 f2f meeting STDT agreed to study/design three mission concepts
- **Concept 1 (science driven)**: Off-axis 9m aperture with five instruments, SLS as the launch vehicle due to mass limits to L2 - not telescope size limits.
- **Concept 2 (science-per-dollar maximized)**: smaller aperture size from the Concept A, based on a number of items, still TBD. Launchable in current 5m fairing sizes. STDT will discuss what these are at the September f2f.
- **Concept 3 (led at JPL; cost-max constrained)**: Maximum mission that can be designed for a total cost of \$3B, keeping mid-IR and far-IR (*request from STDT to JPL to carry out this study at the June f2f*).

*(Study Center has proposed to merge Concept 2 and Concept 3 in a joint GSFC/JPL study; TBD at September f2f)*

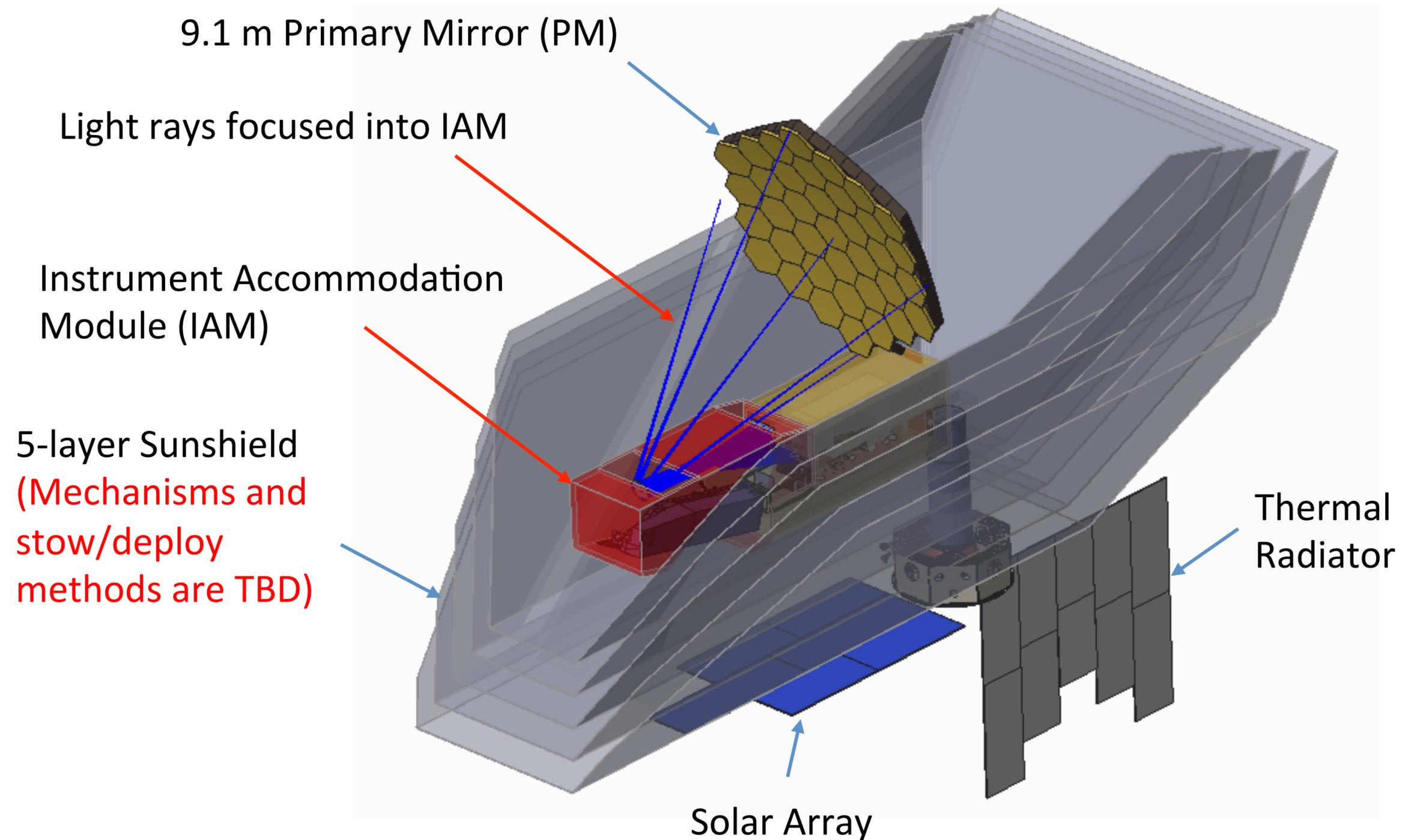
## Top-Level Mission Parameters

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- Mission Duration: 5 years nominal with up to 10 years extended
- Mission Classification: Class A, fully redundant and cross strapped
- Launch Vehicle (LV): Vulcan Aces, Falcon Heavy or similar LV capability
- LV Fairing Size: 5-m Atlas V fairing or similar
- Optical Telescope Element (OTE) Temperature: 4 K *Actively cooled using cryocoolers*
- 9.1-m Aperture Primary Telescope Diameter
- Wavelength Coverage: 6 to 600 microns
- Number of instruments: five (5)
  - Medium Resolution Survey Spectrometer (MRRS) – JPL
  - Hi Res (Far-IR) Spectrometer (HRS) - GSFC
  - Heterodyne Instrument (HI) – Europe
  - FIR Imager/ Polarimeter (FIP) – GSFC
  - MID-IR Imager Spectrometer/ Coronagraph (MISC) – JAXA

# OST Concept 1

## Deployed Configuration



OST Concept 1 Instrument Specifications						
Instrument	Wavelength Coverage	Spectral Resolving Power ( $\lambda/\Delta\lambda$ )	Number of spatial pixels or sky beams	Typical Required Sensitivity:	Other	Lead institution
Mid-Infrared imager/spectrograph/coronagraph (MISC)	6 to 38 $\mu\text{m}$	imager: $R\sim 5-10$ ; IFU: $R>1500$	3'x3' FoV; $\sim 10^7$	photometric: 1 $\mu\text{Jy}$ @10 $\mu\text{m}$	coronagraph $10^{-6}-10^{-7}$ IWA= $1\lambda/D$	JAXA/Japan
Far-Infrared Imager + Polarimeter (FIP)	35 to 240 $\mu\text{m}$ (4 channels)	$R\sim 3.3$	$\sim 500,000$	1 $\mu\text{Jy}$ - 10 mJy (confusion limit)	polarimetry, spectral line filters	GSFC
Mid-Res Survey Spectrometer (MRSS)	30 to 660 $\mu\text{m}$	low-res $\sim 500$ high-res $\sim 10^4$	100 per channel	$10^{-21}$ W/m <sup>2</sup> (spectral line)	6 channels simultaneous	JPL
Heterodyne Receiver (HERO)	66 to 610 $\mu\text{m}$ (7 discrete channels)	$\sim 10^7$	1 - 15	2 mK in 0.2 km/s @ 1 THz	background limited	ESA/CNES
High-Res Spectrometer (HRS)	25 to 200 $\mu\text{m}$	low-res $\sim 8 \times 10^4$ high-res $\sim 5 \times 10^5$	18"x3.6" grating FoV at 160 $\mu\text{m}$	$10^{-21}$ W/m <sup>2</sup> 5 $\sigma$ (spectral line)	photo-counting	GSFC



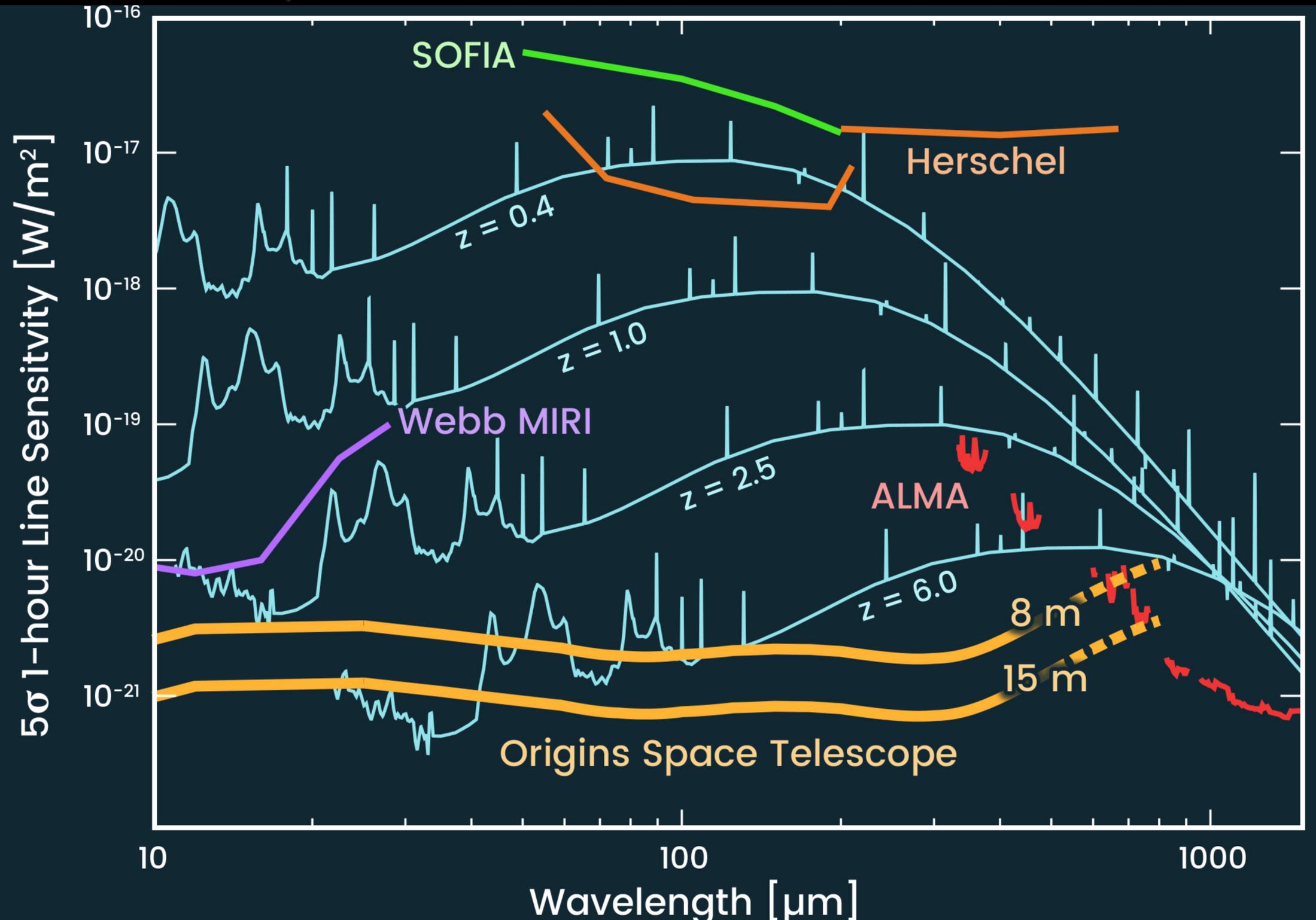
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Two Orders of Magnitude!

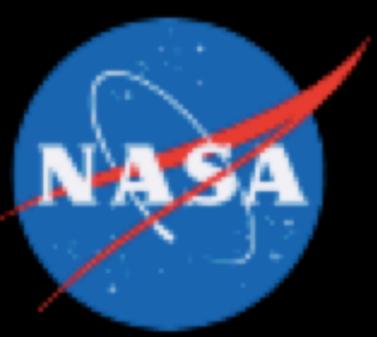



Four Orders of Magnitude!





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## What Origins Space Telescope will be

- A flagship general observatory - community driven sciences and instruments.
- Study gas cloud cooling at cosmic dark ages, to ozone and methane biosignatures of exoplanets, to pathway of water to habitable exoplanets and our Solar system.
- Provides a factor of 10,000 (!) improvement in sensitivity. An immense discovery potential.
- Origins Space Telescope will not be extending what we know already. It will be a true revolution in astronomy.

**@NASAOriginsTelescope**

**[origins.ipac.caltech.edu](http://origins.ipac.caltech.edu)**