



Astrophysics Program Analysis Groups AAS 223rd Meeting Washington, DC

January 5, 2014

Astrophysics

Paul Hertz

**Director, Astrophysics Division
Science Mission Directorate**

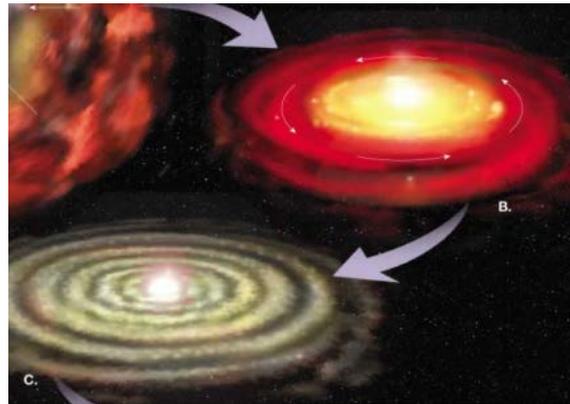


Why Astrophysics?

Astrophysics is humankind's scientific endeavor to understand the universe and our place in it.



1. How did our universe begin and evolve?

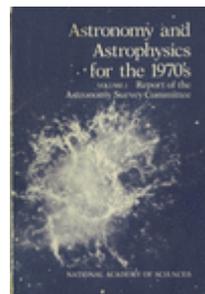


2. How did galaxies, stars, and planets come to be?

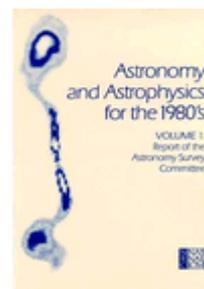


3. Are We Alone?

These national strategic drivers are enduring



1972



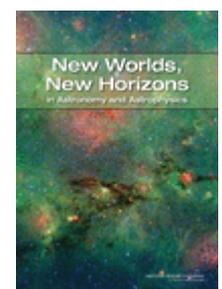
1982



1991



2001

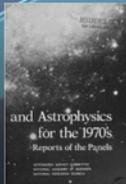


2010

ASTROPHYSICS

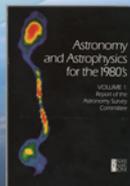
Decadal Survey Missions

1990



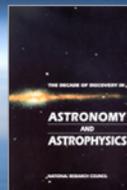
1972
Decadal Survey
Hubble

1999



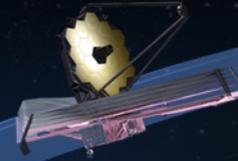
1982
Decadal Survey
Chandra

2003



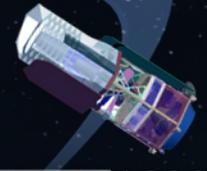
1991
Decadal Survey
Spitzer

LRD: 2018



2001
Decadal Survey
JWST

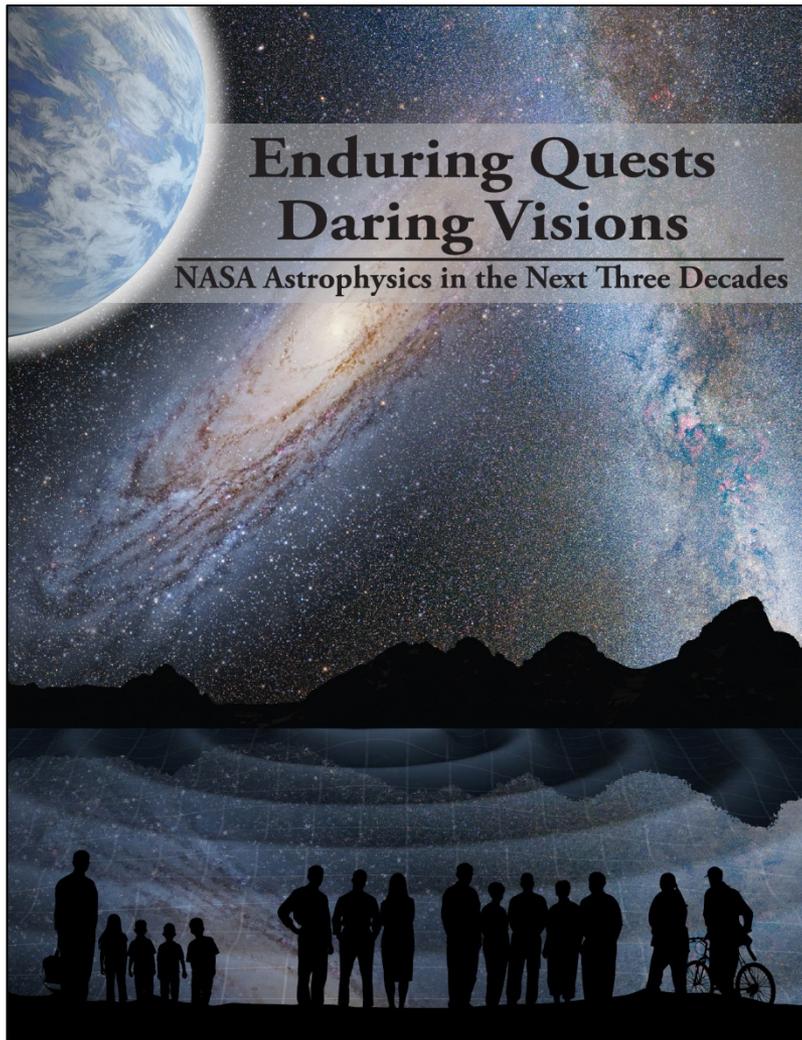
LRD: 2020s



2010
Decadal Survey
WFIRST



Enduring Quests, Daring Visions



- A 30 year vision to address the enduring questions:
 - Are we alone?
 - How did we get here?
 - How does the universe work?

	Near-Term	Formative	Visionary
Gravitational Waves		 Gravitational Wave Surveyor	 Gravitational Wave Mapper
Cosmic rays	 JEM-EUSO		
Radio			 Cosmic Dawn Mapper
Microwaves		 CMB Polarization Surveyor	
Infrared	 JWST	 Far IR Surveyor	
Optical	 WFIRST-AFTA	 LUVOIR Surveyor	 ExoEarth Mapper
Ultraviolet	 Euclid		
X-rays	 TESS	 Gaia	
Gamma rays	 NICER	 Astro-H	 Xray Surveyor
			 Black Hole Mapper

<http://science.nasa.gov/astrophysics/documents>



Community Participation

COPAG

- Chair: Ken Sembach
- Executive Cmte: 10 members
- SAGs: 6 Active
- Website:

<http://cor.gsfc.nasa.gov/copag>

ExoPAG

- Chair: Scott Gaudi
- Executive Cmte: 11 members
- SAGs: 3 Active
- Website:

<http://exep.jpl.nasa.gov/exopag>

PhysPAG

- Chair: John Nousek
- Executive Cmte: 13 members
- SAGs: 5 Active
- Website:

<http://pcos.gsfc.nasa.gov/physpag>

Science and Technology Definition Teams (STDTs):

- AFTA use of telescope assets: 20 members
- Exoplanet Probe with Internal Coronagraph: 10 members
- Exoplanet Probe with External Occulter: 10 members
- X-ray Astrophysics Probe: 14 members [disbanded 12/12/13]

Preliminary reports from the studies are due Spring 2014.

Final reports from the studies are due in January 2015.

Advisory Committees:

- NRC Space Studies Board (SSB)
- NRC Committee on Astronomy and Astrophysics (CAA)
- Astronomy and Astrophysics Advisory Committee (AAAC)
- NASA Advisory Council's Science Committee (NAC SC)
- NASA Advisory Council's Astrophysics Subcommittee (APS)



PAG Structure

NASA Advisory Council (NAC)

Science Committee

Astrophysics Subcommittee

COPAG

ExoPAG

PhysPAG

Chair: Ken Sembach
Executive Cmte: 10 members
SAGs: 6 Active
<http://cor.gsfc.nasa.gov/copag>

Chair: Scott Gaudi
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<http://exep.jpl.nasa.gov/exopag>

Chair: John Nousek
Executive Cmte: 13 members
SAGs: 5 Active
<http://pcos.gsfc.nasa.gov/physpag>

- COPAG SAGs/SIGs include
- Science objectives for a 4m–8m UV/optical mission (closed)
 - Technologies for a 4m-class monolithic telescope UV/optical mission w/internal coronagraph (closed)
 - Technologies for an 8m-class segmented telescope UV/optical mission w/external occulter (closed)
 - Technologies for a future far-IR mission
 - Science objectives & technology requirements for a series of Cosmic Origins Probes
 - Cosmic origins science enabled by the WFIRST-AFTA coronagraph
 - Science enabled by operations overlap of HST and JWST
 - Science enabled by the WFIRST - AFTA data archive
 - Far infrared science and technology Science Interest Group

- ExoPAG SAGs include
- Debris Disks & Exozodiacal Dust (closed)
 - Potential for Exoplanet Science Measurements from Solar System Probes (closed)
 - Planetary Measurements Needed for Exoplanet Characterization
 - Exoplanet Flagship Requirements and Characteristics (closed)
 - Requirements and Limits of Future Precision Radial Velocity Measurements
 - Exoplanet Probe to Medium Scale Direct-Imaging Mission Requirements and Characteristics
 - Characterizing the Atmospheres of Transiting Planets with JWST and Beyond
 - Preparing for the WFIRST Microlensing Survey

- PhysPAG SAGs/SIGs include:
- Technology SAG (closed)
 - Inflation Probe Science Interest Group
 - Gravitational Wave Science Interest Group
 - X-ray Science Interest Group
 - Gamma-ray Science Interest Group
 - Cosmic Ray Science Interest Group



Present

Near Term

Formative

Visionary

Science Roadmap

Missions

	Present	Measure dark energy & history of cosmic growth		Map structure at reionization
		Probe the epoch of inflation		Measure cosmic expansion history with standard sirens
		Completely characterize the CMB		
	Present	Constrain neutron star equation of state		Map black holes using gravitational waves
		Understand black-hole-powered engines		Measure black hole masses & spins
		Image sources detected by aLIGO		Image the shadows of black hole event horizons
				Chart supermassive black hole mergers
				Search for electroweak-era gravitational waves
				Hear the Big Bang

Science Roadmap

Missions

	Discover
	Map the
	Find the
	Image

Hubble	LSST	Gravitational Wave Surveyor	Gravitational Wave Mapper
Spitzer	Extremely Large Telescopes	X-ray Surveyor	X-ray Mapper
Herschel	James Webb Space Telescope	LUVOIR Surveyor	ExoEarth Mapper
ALMA	WFIRST-AFTA	Far Infrared Surveyor	Cosmic Dawn Mapper

Science Roadmap

	Complete
	Characterize
	Measure potential

Missions

Kepler	TESS	LUVOIR Surveyor	ExoEarth Mapper
Hubble	James Webb Space Telescope		
Spitzer	WFIRST-AFTA		

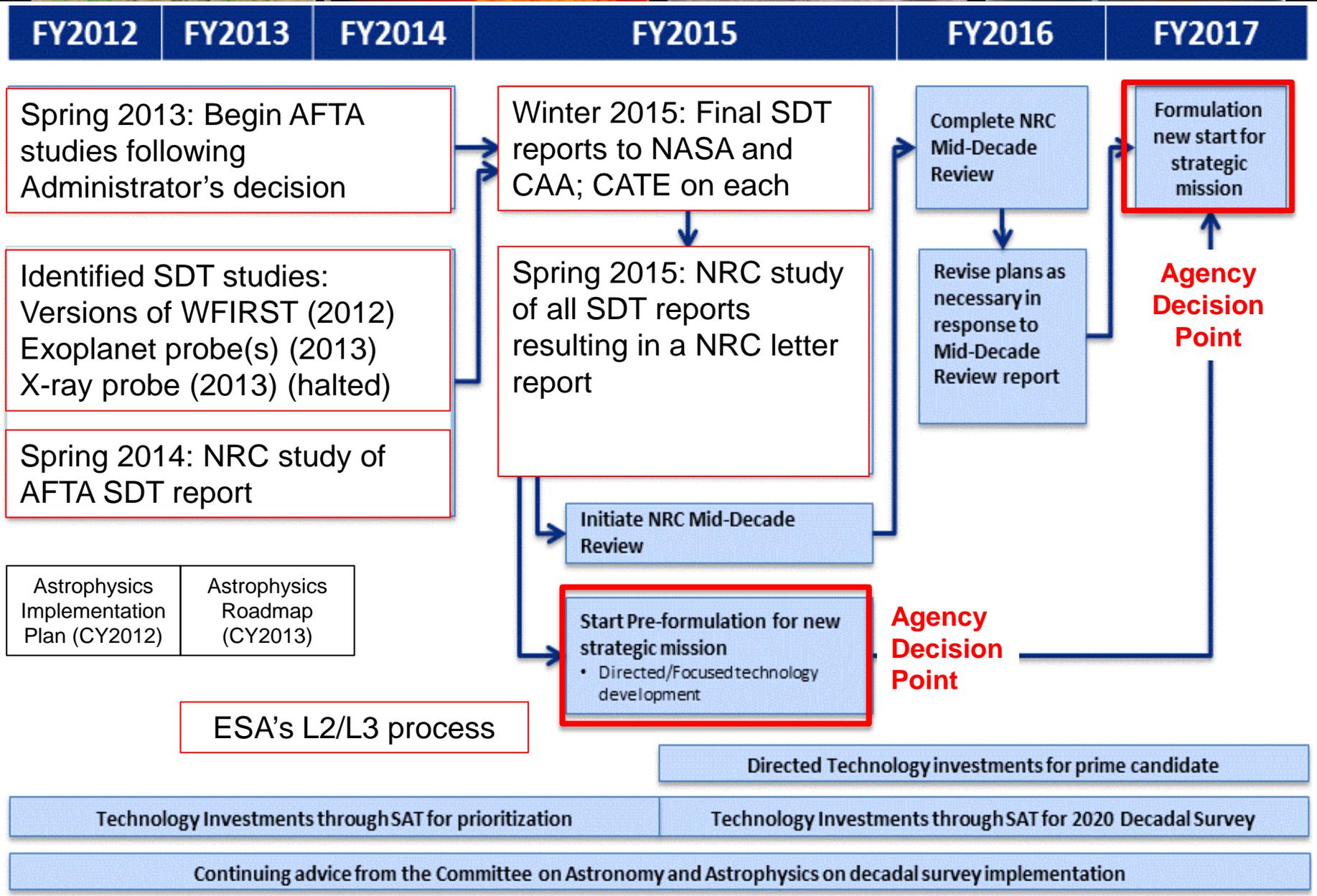


Role of PAGs

- A PAG enables direct regular communication with the community through public meetings that give the community opportunities to provide its scientific and programmatic input.
 - A PAG provides findings of analyses to NASA through the NASA Advisory Council within which the PAG Chair is a member of the Astrophysics Subcommittee.
- A PAG is responsible for soliciting and coordinating community input into the **Tuesday 2:00 – 3:30**
 - The **Session 234: Reports from NASA's Astrophysics** members input who **Program Analysis Groups** input with **(Potomac Ballroom A)**
 - The full PAG consists of all members of the community who participate in the PAG's open meetings.
- The PAG may choose to organize sub-groups (e.g., Science Analysis Groups - SAGs and/or Science Interest Groups- SIGs) to deal with specific issues and report their findings to the full group.
 - A SAG is typically tasked with reporting on a specific issue, and when the analysis of that issue is complete the SAG is disbanded.
 - A SIG is typically tasked with collecting community input from a specific community on a longer-term basis.



Implementing the Decadal Survey





Progress Toward Decadal Survey Priorities

The President's Budget Request for FY14 supports

L1. WFIRST	Preformulation and focused technology development for AFTA (a 2.4m version of WFIRST) are underway to preserve a new start NET FY17
L2. Augmentation to Explorer Program	Increased from ~\$90M in FY07 and ~\$115M/yr in FY10 to ~\$140M/yr in FY16 and beyond; supports AOs in 2014, 2017, ...
L3. LISA	Strategic technology investments plus discussing partnership in ESA's L3 gravitational wave observatory
L4. IXO	Strategic technology investments plus discussing partnership in ESA's L2 X-ray observatory
M1. New Worlds Technology Development Program	Focused technology development for a coronagraph on WFIRST; probe mission concept studies and strategic technology investments
M2. Inflation Probe Technology Development Program	Three balloon-borne investigations plus strategic technology investments

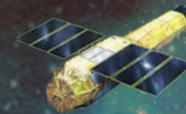
- Formulation
- Implementation
- Primary Ops
- Extended Ops



XMM-Newton (ESA)
12/10/1999



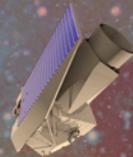
Swift
11/20/2004



Suzaku (JAXA)
7/10/2005



Fermi
6/11/2008



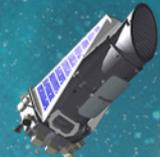
Euclid (ESA)
2020



Spitzer
8/25/2003



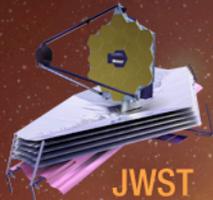
Hubble
4/24/1990



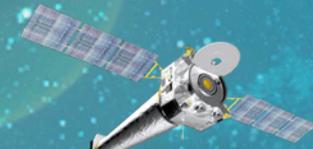
Kepler
3/6/2009



Astro-H (JAXA)
2015



JWST
2018

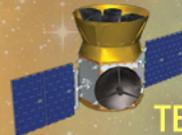
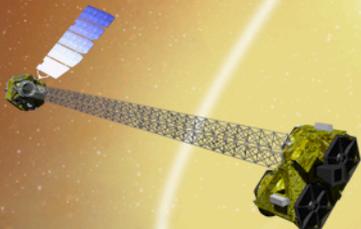


Chandra
7/23/1999



NICER (on ISS)
2016

NuSTAR
6/13/2012



TESS
2017



LISA Pathfinder (ESA)
2015



SOFIA
Full Ops 2014

NASA Astrophysics Programs

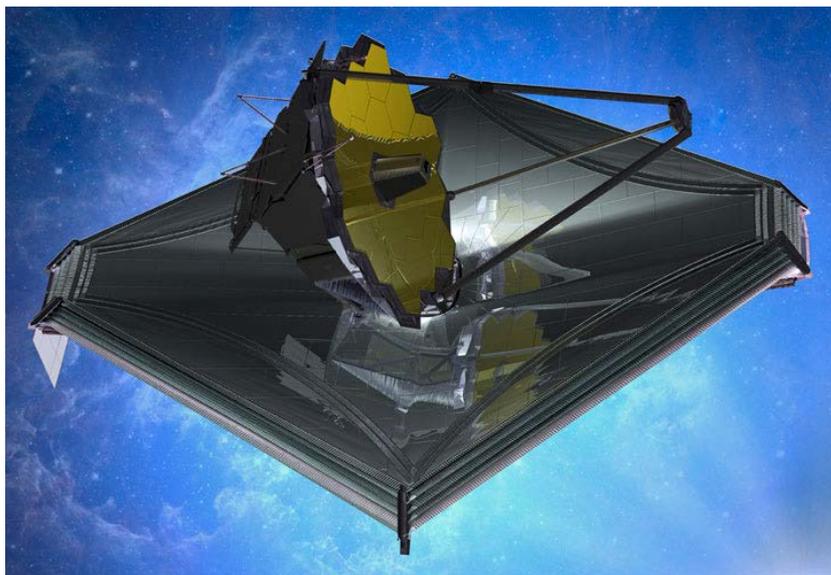
- Recently Completed

 - Planck 2013
 - Herschel 2013
 - GALEX 2013



JWST

James Webb Space Telescope



Large Infrared Space Observatory

Top priority of 2000 Decadal Survey

Science themes: First Light; Assembly of Galaxies; Birth of Stars and Planetary Systems; Planetary Systems and the Origins of Life

Mission: 6.5m deployable, segmented telescope at L2, passively cooled to <50K behind a large, deployable sunshield

Instruments: Near IR Camera, Near IR Spectrograph, Mid IR Instrument, Near IR Imager and Slitless Spectrograph

Operations: 2018 launch for a 5-year prime mission

Partners: ESA, CSA

CURRENT STATUS:

- Project has entered its long and challenging Integration and Test activities.
- Technical progress continues to be significant.
 - Instruments are delivered and in integration & test phase.
 - All optics are complete (primary segments, secondary, tertiary and fine steering mirrors) and delivered to GSFC.
 - Telescope wings are complete; backplane support fixture and center section are complete.
 - Spacecraft completing reviews leading to spacecraft Critical Design Review (Jan 2014).
- Project is performing within the budget, to schedule.
 - Government shutdown did not impact October 2018 launch date.
- FY14 is the peak funding year with many critical activities.



JWST

James Webb Space Telescope

- JWST remains on schedule for its October 2018 launch

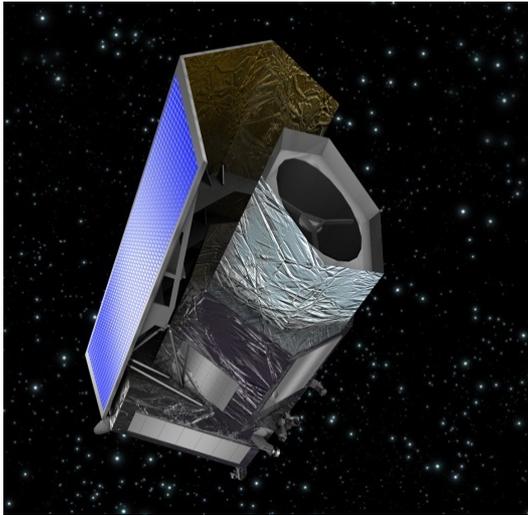
- JWST Town Hall: Wednesday 12:45 pm in Potomac Ballroom A





Euclid

A visible and near-infrared telescope to explore cosmic evolution



CURRENT STATUS:

- Currently in implementation phase.
- ~50 U.S. scientists are members of the Euclid Science Team that will analyze the data, and make maps of the sky.
- First experimental manufacturing run for the Euclid near-infrared detectors to complete in FY 2014 (ESA).
- NASA will initiate the buy for the flight infrared detectors in FY 2014.
- NASA will test and characterize the near-IR flight detectors.

- **ESA Cosmic Vision 2015-2025 Mission,** M-Class with NASA participation.
- 1.2-m mirror, visible & near-IR images, spectra
- **Launch Date:** Mar 2020, 5 year prime mission
- **Science Objectives:**
 - Euclid will look back 10 billion years into cosmic history.
 - Probe the history of cosmic expansion (influenced by dark energy and dark matter) and how gravity pulls galaxies together to form the largest structures.
 - The shapes of distant galaxies appear distorted because the gravity of dark matter bends their light (gravitational lensing). Measuring this distortion tells us how the largest structures were built up over cosmic time.
 - Measuring how strongly galaxies are clumped together tells us how gravity influences their motions, and how dark energy has affected the cosmic expansion.



WFIRST – AFTA

Widefield Infrared Survey Telescope with Astrophysics Focused Telescope Assets



WFIRST Town Hall: Wednesday 6:30 pm
in National Harbor 3

- **Top priority in 2010 Decadal Survey**
- **Study Baseline Payload:**
 - 2.4m existing telescope assets
 - Widefield imager
 - Coronagraph
- **Science objectives:**
 - Hubble-quality imaging over 200x the field
 - Comprehensive study of dark energy
 - Systematic census of outer planets
 - Coronagraphic imaging of exoplanets
 - 25% time for community competitive selected GO program
 - Enhancing JWST science

CURRENT STATUS:

- May 2013, NASA Administrator Bolden directed Astrophysics Division to study WFIRST-AFTA and preserve option for FY17 new start if budget is available
 - No decision expected before early 2016
- Currently in pre-formulation phase
 - NRC study in early 2014
 - SDT final report due Jan 2015
- Maturing key technologies to TRL 5 by FY17 and TRL 6 by FY19
 - Infrared detectors
 - Internal coronagraph for exoplanet characterization

Mission description:

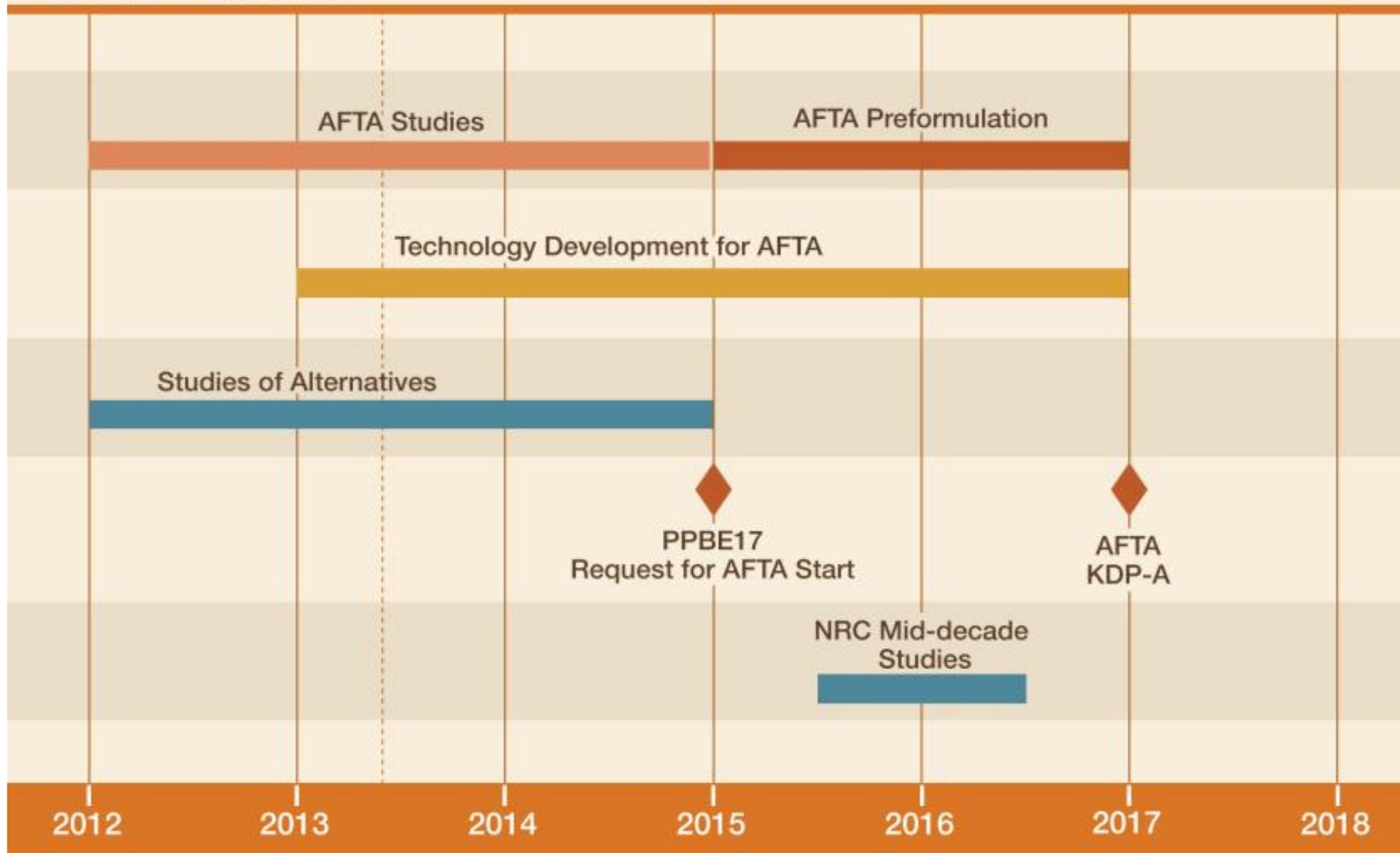
- #1 Large-Scale Priority: Widefield infrared survey telescope for Dark Energy, Exoplanets, IR Surveys
- #1 Medium-Scale Priority: Development and demonstration of technology for direct imaging and characterization of exoplanets



Plan for WFIRST-AFTA Preformulation

Widefield Infrared Survey Telescope using Astrophysics Focused Telescope Assets

AFTA timeline





WFIRST – AFTA Coronagraph

- Community AFTA coronagraph working group (ACWG) conducted an open, technical evaluation using public evaluation criteria in a series of workshops and telecons since July 2013, reaching a broad consensus on the basis for the decision
- Three strong technologies emerged, spanning the risk/performance continuum
- The independent Technical Analysis Committee (TAC) concurred with the basis and with findings of ACWG

Decision:

- **Primary Architecture:** Occulting Mask Coronagraph (OMC) that includes masks for Shaped Pupil Coronagraph (SPC) and Hybrid Lyot Coronagraph (HLC)
- **Backup Architecture:** Phase-Induced Amplitude Apodization Complex Mask Coronagraph (PIAACMC)
- Combination best minimizes risk, preserves options to protect the project schedule, advances technologies, and preserves possibilities of increased science yield
- Plan for Recommendation to reach TRL 5 is feasible (technically) and credible within existing resources (schedule, cost)



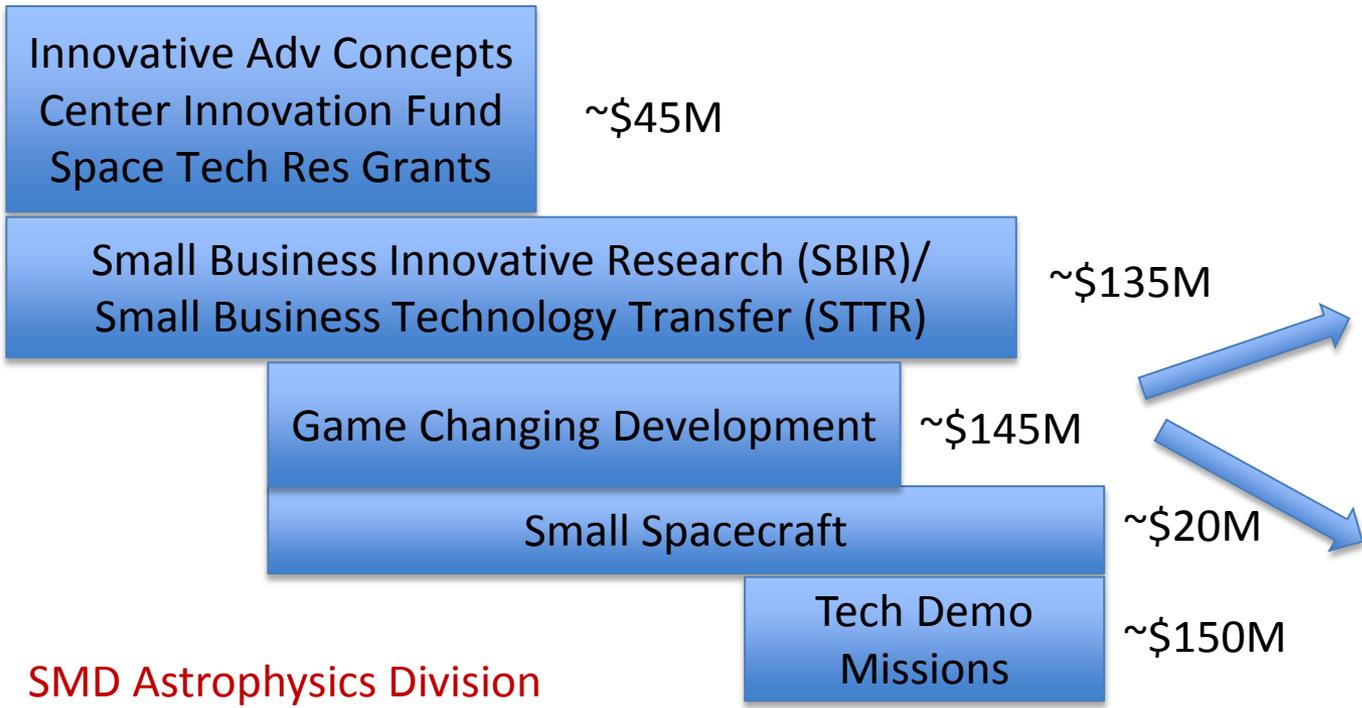
ESA's New Vision to Study the Invisible Universe

- The hot energetic Universe and the search for elusive gravity waves will be the focus of ESA's next two large science missions.
- The science theme “the hot and energetic Universe” was selected for L2, and expected to be pursued with an advanced X-ray observatory.
 - Launch date ~2028
- The L3 mission will study the gravitational Universe, searching for ripples in the very fabric of space-time created by celestial objects with very strong gravity, such as pairs of merging black holes.
 - Launch date ~2034
 - Will require development of a spaceborne gravitational wave observatory or extreme precision ‘gravitometer’.
- NASA has expressed a strong interest to ESA in contributing to ESA's next large astrophysics missions if they are responsive to the U.S.
Decadal Survey
 - The U.S. Decadal Survey recommended an international partnership for a gravitational wave observatory and an X-ray observatory.
- NASA and ESA will have a bilateral meeting, on January 8-10, 2014 in D.C., to discuss a potential NASA contribution.

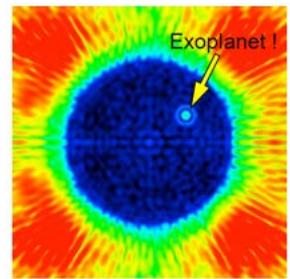
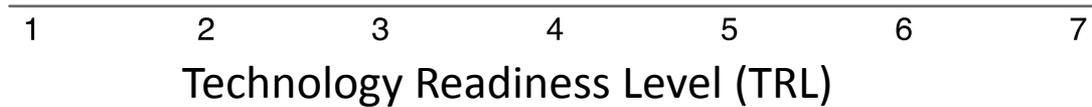
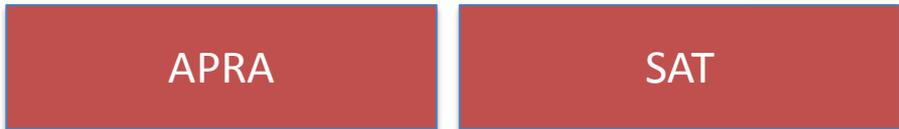


Space Technology Mission Directorate Contributes to Astrophysics

Space Technology Mission Directorate



SMD Astrophysics Division

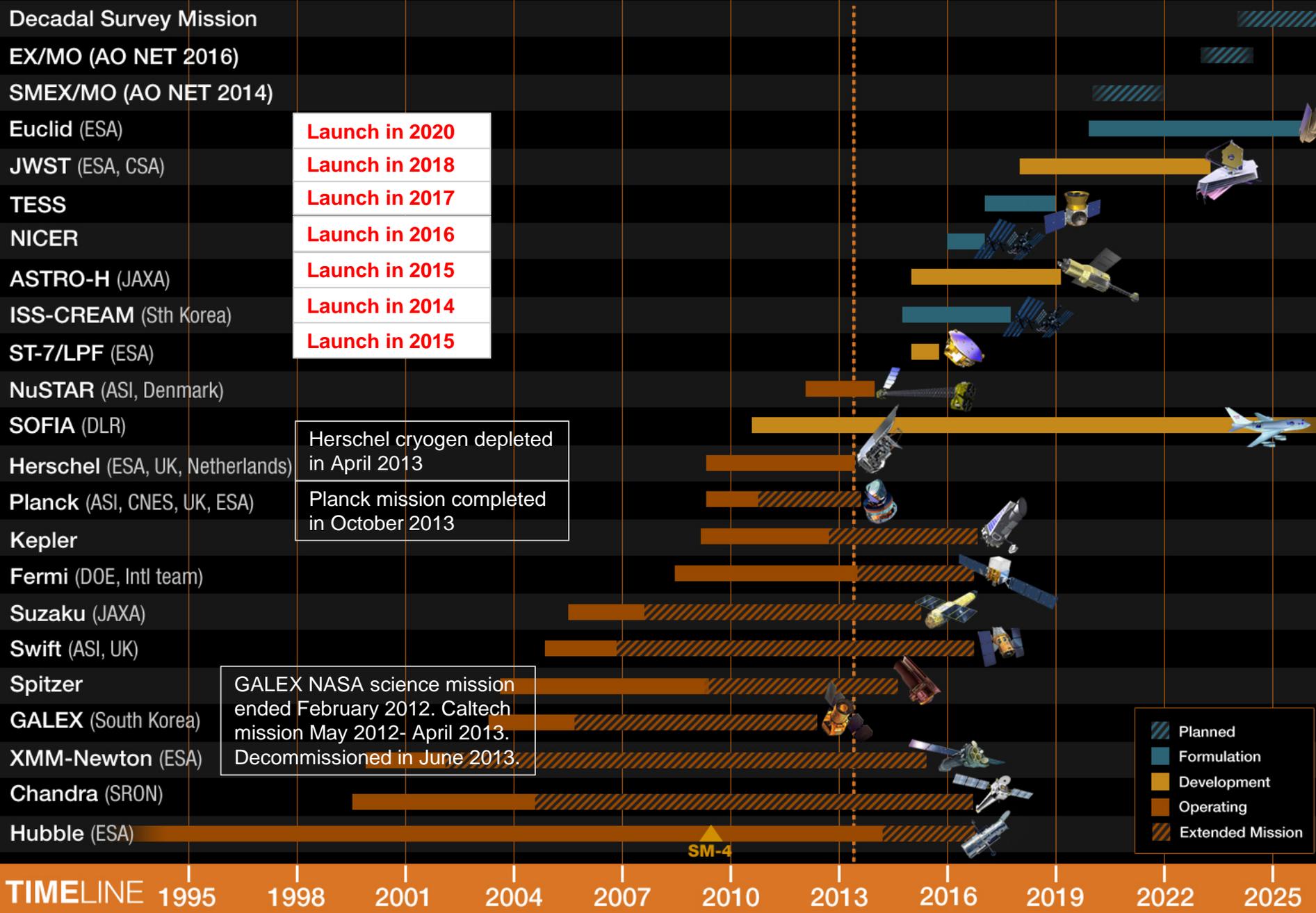




Strategic Astrophysics Technology

- Decadal Survey recommended funding for medium-term technology development be augmented.
- The SAT Program is designed to support the maturation of technologies from demonstrated feasibility (i.e., TRL 3) to where they can be incorporated into NASA flight missions (TRL 6-7).
- SAT has several components
 - Directed technology development (e.g. high contrast imaging testbeds)
 - Focus on **Wednesday 6:30 – 8:00** /)
 - **Session 339: Preparing for Future NASA Missions:**
- SAT first **The Strategic Astrophysics Technology Program**
 - **Start (National Harbor 2)**
 - More than 35 selected technology investigations to date
 - Some years impacted by changes in the astrophysics planning budget or sequestration
 - Balance between competed and directed has evolved with commitment to having AFTA ready for a formal start by FY2017
- Priorities given to technologies required for future strategic missions
 - WFIRST, GW observatory, X-ray observatory, exoplanet characterization, CMB polarization, ...
- SAT will be fully offered in ROSES-14 with ~\$6M in new awards

Astrophysics Missions timeline



Launch in 2020
Launch in 2018
Launch in 2017
Launch in 2016
Launch in 2015
Launch in 2014
Launch in 2015

Herschel cryogen depleted in April 2013
 Planck mission completed in October 2013

GALEX NASA science mission ended February 2012. Caltech mission May 2012- April 2013. Decommissioned in June 2013.

Planned
 Formulation
 Development
 Operating
 Extended Mission

SM-4

TIMELINE 1995 1998 2001 2004 2007 2010 2013 2016 2019 2022 2025

Monday

12:45 – 1:45 Session 122: Kepler Mission Town Hall (Potomac Ballroom C)

6:30 – 8:00 SOFIA Mission Update (Maryland Ballroom A)

Tuesday

12:45 – 1:45 Session 222: NASA Town Hall (Potomac Ballroom A)

2:00 – 3:30 Session 234: Reports from NASA's Astrophysics Program Analysis Groups (Potomac Ballroom A)

6:30 – 8:00 Exoplanet Exploration Program Town Hall (National Harbor 3)

Wednesday

12:45 – 1:45 Session 319: Hubble and James Webb Space Telescope Town Hall (Potomac Ballroom A)

6:30 – 8:00 Session 339: Preparing for Future NASA Missions: The Strategic Astrophysics Technology Program (National Harbor 2)

6:30 – 8:00 Session 341: Wide Field Infrared Survey Telescope (WFIRST) Town Hall (National Harbor 3)

Thursday

10:00 – 11:30 Session 414: Science Highlights from NASA's Completed Astrophysics Data Analysis Program (Potomac Ballroom A)

■ Formulation
■ Implementation
■ Primary Ops
■ Extended Ops

NuSTAR
6/13/2010

IWCT
2018

XMM-Newton (ESA)
10/10/1999

Swift
11/13/2004

Suzaku (JAXA)
1/10/2005

6/11/2008

Spitzer
2003

2000

Hubble
4/24/1990

Astro-H (JAXA)
2015

Kepler
3/6/2009

Chandra
7/23/1999

ICER (NASA)
2010

2017

Completed
Planck 2013
Herschel 2013
GALEX 2013

SOFIA
Full Ops 2014



Backup



The Big Picture

- This remains a time of scientific opportunity for NASA Astrophysics.
 - We are poised to answer the most compelling science questions.
 - The budget for NASA astrophysics, which includes JWST, is at ~\$1.25B, a high level.
 - NASA continues to operate large and small space-based observatories spanning the electromagnetic spectrum, including multiple Great Observatories.
 - The James Webb Space Telescope, the highest priority of the community, is on schedule and fully funded for an October 2018 launch.
 - NASA continues to develop contributions to international missions for launch this decade.
 - NASA has downselected two new Explorer projects to begin development for launch this decade, and an Explorer AO is planned for late 2014 to select two more Explorer projects.
 - NASA continues to support individual investigators for data analysis, theory, and technology investigations through open, competitive, peer reviewed processes.
 - NASA is preparing for the strategic mission that will follow JWST.
- The budgetary future remains uncertain.
 - Priorities must be used to guide difficult budget choices.



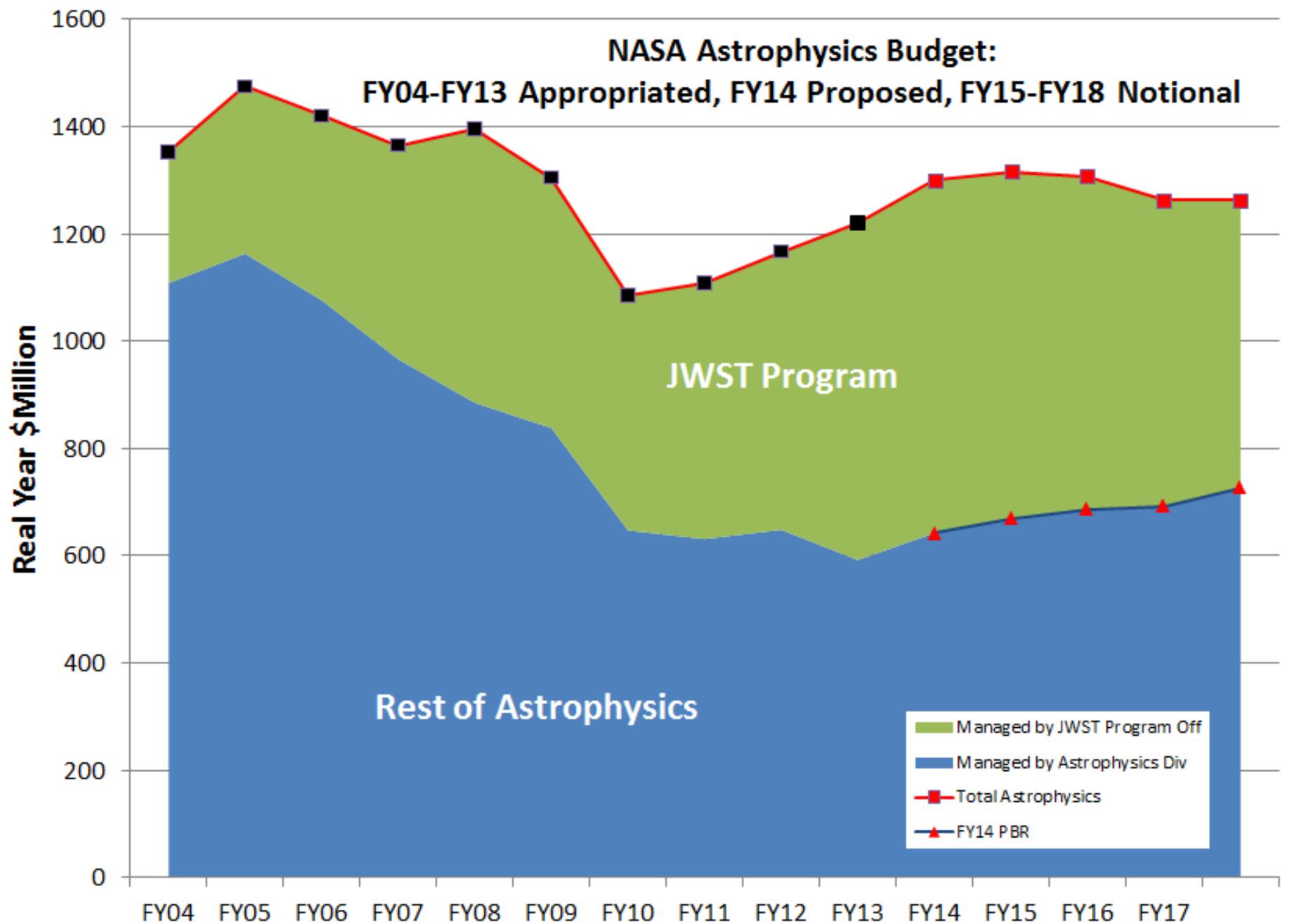
Astrophysics Decadal Survey - Summary

Program Scale	Recommendation	Response supported by FY14 President's Budget Request
Large	WFIRST	DRM1 and DRM2 completed in FY12; AFTA "proof of concept" DRM completed in FY13; pre-formulation and technology development (detector and coronagraph) in FY14-FY19; prepared for decision regarding new start in FY15; participating in ESA's Euclid
Large	Explorer Augmentation	Impacted by sequestration and budget reductions including cancellation of selections from FY12 MO AO; EX AO in FY11; SMEX AO NET 2014; EX AO NET 2016; each AO has a mission and a MO
Large	LISA Technology	CST completed in FY12; technology supported through SAT; ST-7/LPF supported; will pursue partnership with ESA if a GW mission is selected for L2/L3 mission
Large	IXO Technology	CST completed in FY12; technology supported through SAT; X-ray probe STDT starting in FY14; will pursue partnership with ESA if an X-ray mission is selected for L2/L3 mission
Medium	New Worlds Technology	Technology supported through APRA and SAT(TDEM); exoplanet probe STDTs started in FY13; AFTA coronagraph study completed in FY13; AFTA coronagraph technology starting in FY14; will consider partnership with ESA if an exoplanet mission is selected for L2/L3 mission
Medium	Inflation Probe Technology	Technology supported through APRA and SAT including multiple suborbital payloads; will consider partnership with ESA if a CMB mission is selected for L2/L3 mission
Small	Astrophysics Theory Program Augmentation	Impacted by sequestration and budget reductions
Small	(Definition of) a future UV-optical space capability	RFI in FY12; follow-on workshops FY14-FY16; technology supported through APRA, SAT, and working with STMD
Small	Intermediate Technology Development Augmentation	SAT program initiated in FY11 and funded for prioritized investments; funding directed toward decadal survey priorities including AFTA, probes, New Worlds, and ESA L2/L3 technologies; impacted by sequestration and budget reductions
Small	Laboratory Astrophysics Augmentation	Augmentation started in FY12 including selection of large consortium; future selections impacted by sequestration and budget reductions
Small	SPICA mission (U.S. contributions to JAXA-led)	Candidate for future Explorer Mission of Opportunity
Small	Suborbital Program Augmentation	Technology augmentation for balloon program; continued development of ULDB balloon platforms; ISS payload selections; impacted by sequestration and budget reductions
Small	Theory and Computation Networks (NASA, NSF, DOE)	Six networks competitively selected in 2013 and funded by NSF and NASA in FY14-FY16
N/A	Additional core program augmentations	Includes basic research and technology development, mission extensions, data analysis, N.G. Roman Technology Fellowships; impacted by sequestration and budget reductions



Astrophysics Budget Strategy

- Use the scientific priorities of the 2010 Decadal Survey to guide strategy and inform choices.
- There is inadequate available budget to implement the 2010 Decadal Survey recommendations as written.
- A goal is to be prepared to start a new strategic NASA Astrophysics mission to follow JWST as soon as funding becomes available, while continuing to advance Decadal Survey science during the interim.
 - WFIRST-AFTA (WFIRST using existing 2.4 m telescopes)
 - Moderate missions (“probes”) derived from the science objectives of the prioritized missions and recommendations in the 2010 Decadal Survey are being studied, in addition to a large mission (WFIRST), to be prepared for a mid-decade decision.
- As appropriate, collaborate with international partners to realize Decadal Survey priorities and recommendations.
 - Partner on ESA’s Euclid mission (complements WFIRST commitment)
 - Partner on ESA’s L2 x-ray observatory (responds to IXO recommendation)
 - Partner on ESA’s L3 gravitational wave observatory (responds to LISA recommendation)





Distribution of FY14 Budget Request

	% of FY14 PBR	Total \$628.4M (excludes \$13.9M SMD admin account)
R&A program elements	13.2%	includes APRA, OSS, ATP, ADAP, RTF, TCAN
Research infrastructure	10.2%	includes balloon program, Keck, LBTI, archives, astrobiology
Einstein, Hubble, Sagan Fellowships	2.2%	
Operating missions (including GO programs)	<u>Total 36.2%</u> Hubble 15.3% Chandra 8.7% Kepler 3.0% Spitzer 2.6% Fermi 2.3% Others 4.4%	prioritized by Senior Review “others” includes Herschel, NuSTAR, Planck, Swift, Suzaku, XMM-Newton GO funding is 9.6% (of \$628.4M FY14 PBR)
SOFIA	13.9%	
Explorer missions in development	12.8%	includes ASTRO-H, NICER, TESS
Strategic missions in development	2.9%	includes Euclid, ST-7
Future Explorer missions	0.0%	no funding until next AO selection
Pre-formulation of WFIRST/AFTA	2.1%	including technology development for detectors and coronagraph
Strategic Astrophysics Technology	3.3%	directed, competed, and testbeds
Other strategic studies	0.7%	includes exoplanet probes, X-ray probe
Program management	2.6%	



AFTA Study: Strawman Payload & SDT Findings

2.4m Telescope with wide field-of-view

Wide-Field Instrument

- *Imaging & spectroscopy over 1000s sq deg.*
- *Monitoring of SNe and microlensing fields*
- 0.7 – 2.0 micron bandpass
- 0.28 sq deg FoV (100x JWST FoV)
- 4 filter imaging, grism + IFU spectroscopy
- 18 H4RG detectors (288 Mpixels)

Requires focused tech. development

Coronagraph (study option)

- *Imaging of ice & gas giant exoplanets*
- *Imaging of debris disks*
- 400 – 1000 nm bandpass
- 10^{-9} contrast
- 100 milliarcsec inner working angle at 400 nm

Requires focused tech. development

Findings of SDT

- AFTA carries out the WFIRST science program (the top ranked decadal priority).
- AFTA's larger aperture enables astronomers to make important contributions towards many of the enduring questions listed in the decadal survey through both surveys and peer-reviewed observing programs.
- Equipped with a coronagraph, AFTA can image Jupiter and Saturn-like planets around the nearest stars. AFTA will be an essential stepping stone towards finding signs of life around nearby stars.

