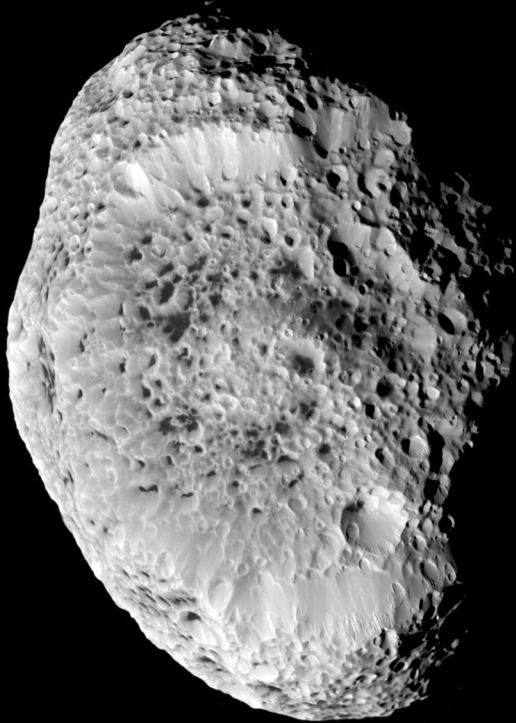


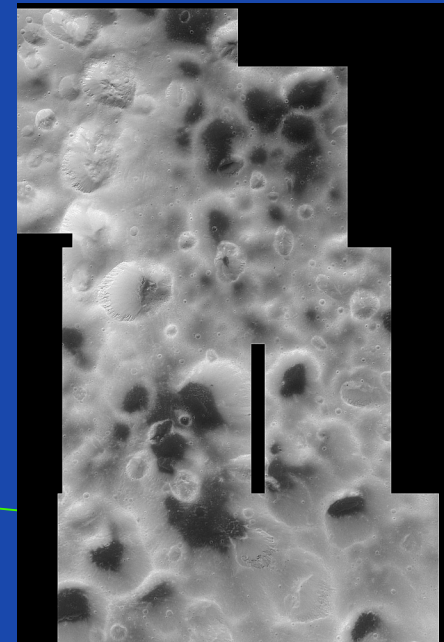
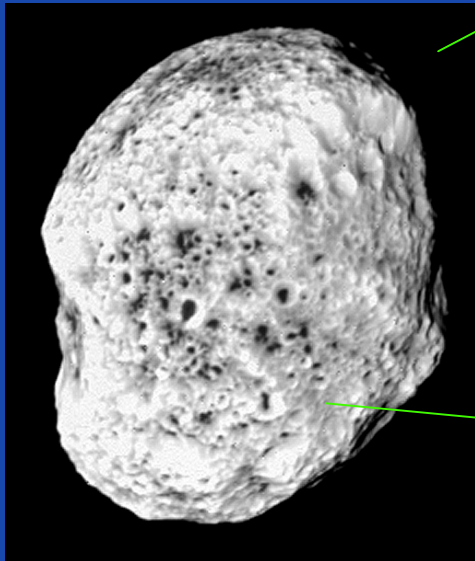
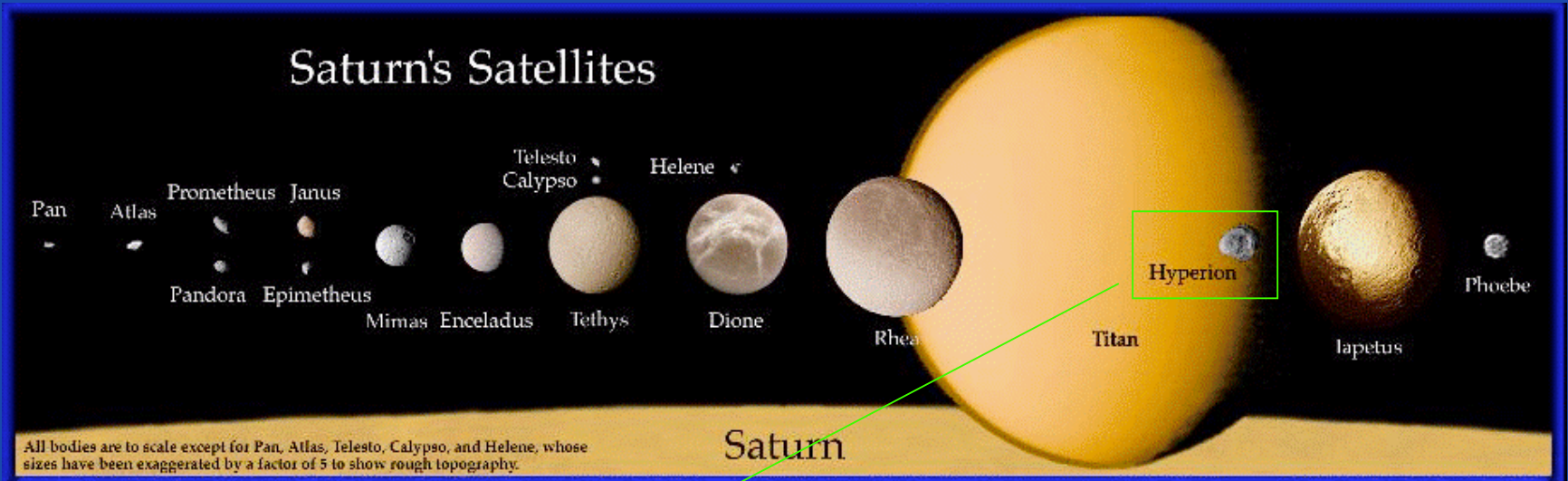
HYPERION: The Sponge Moon



Dr. James Bauer, JPL
Dr. Peter Thomas, Cornell University

August 28, 2007

Hyperion & the Saturnian System

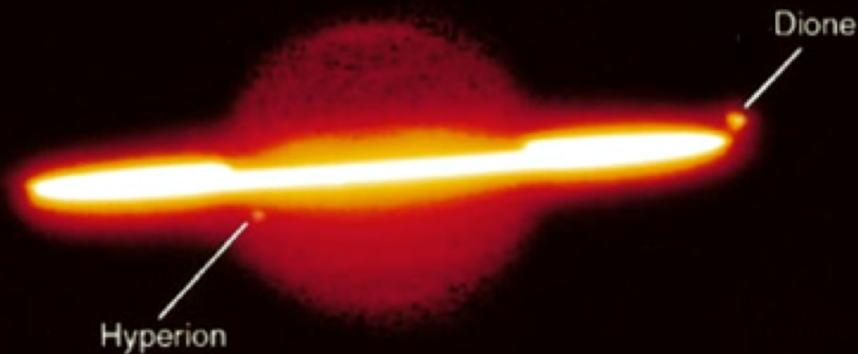


Hyperion from the Ground

Bauer et al. 97

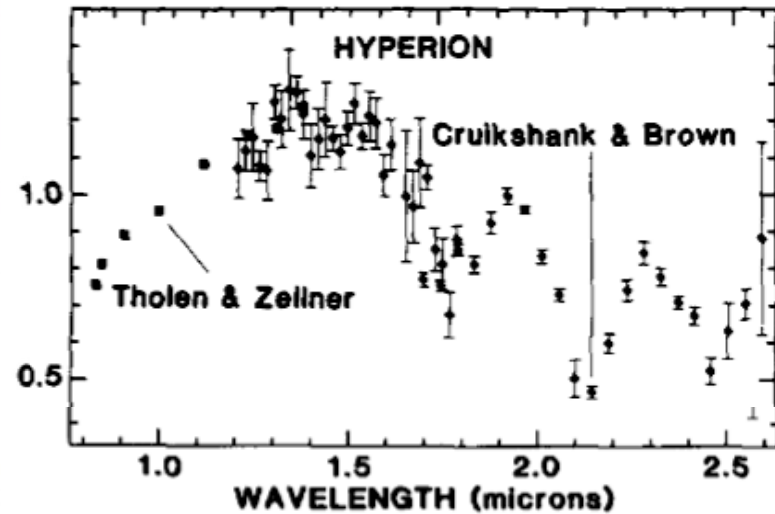


Saturn RPX



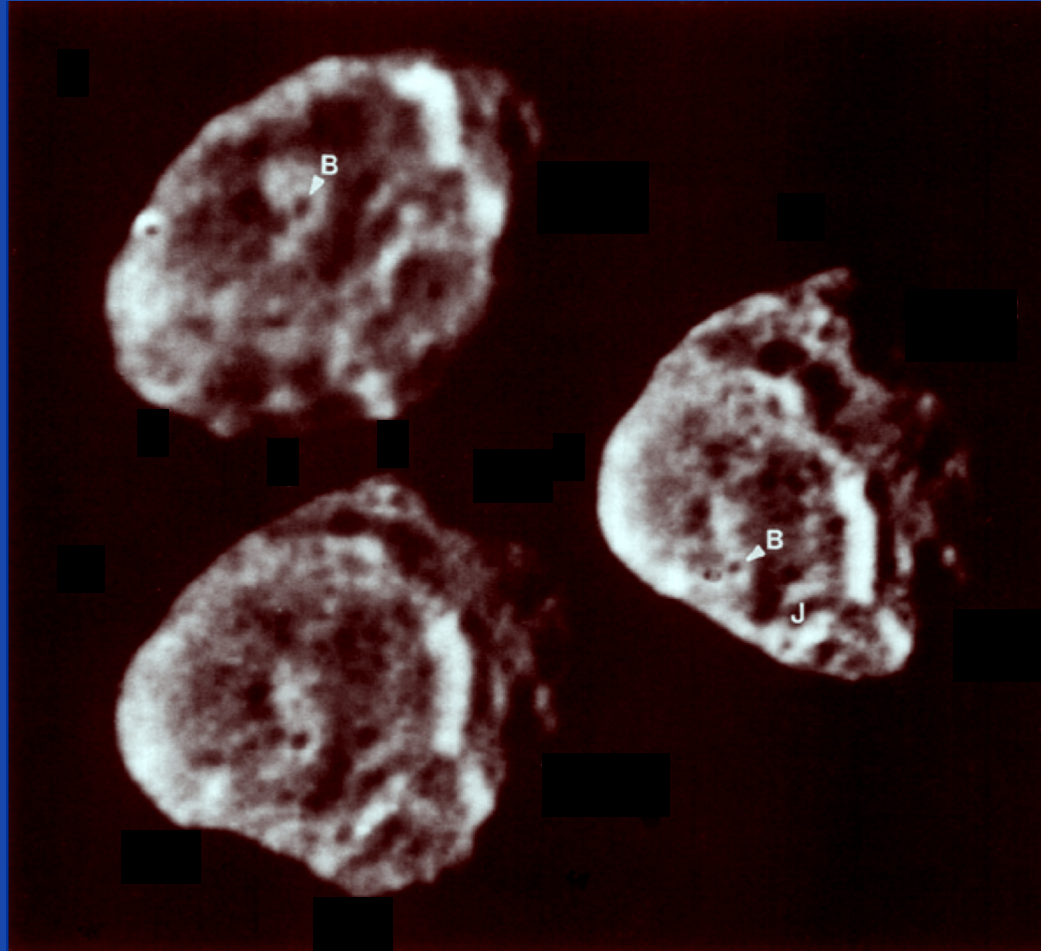
Mommery et al. 2000

September 16, 1995



R. H. Brown 85

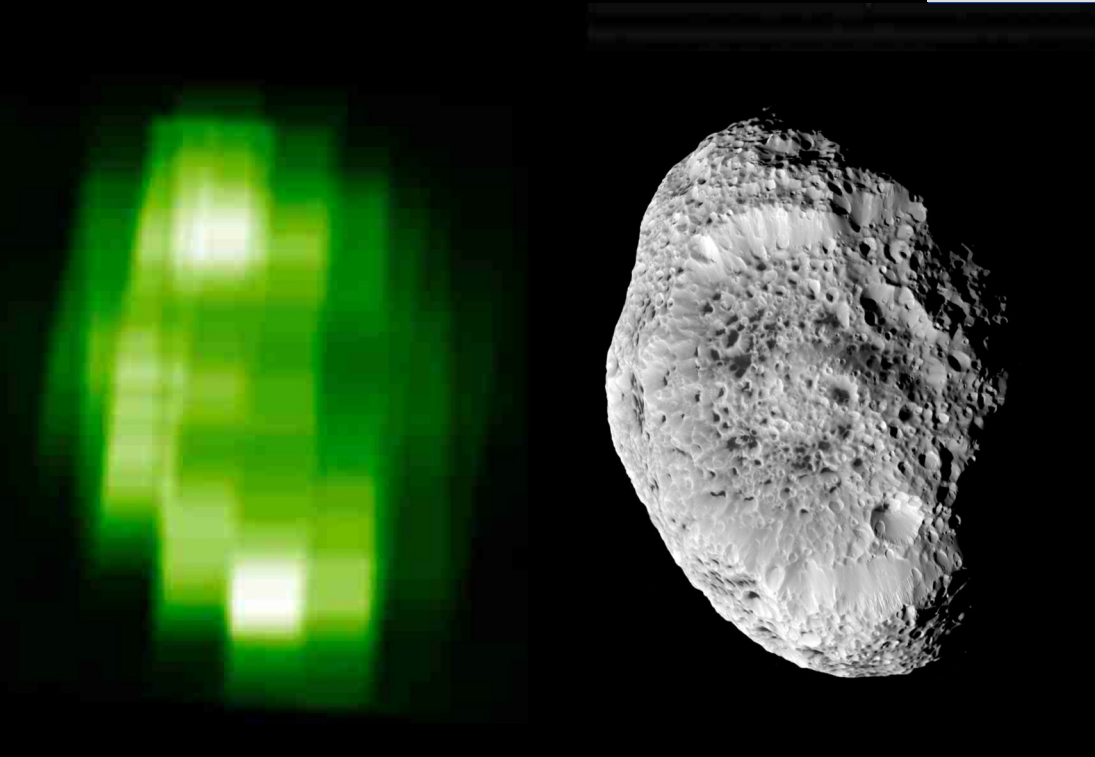
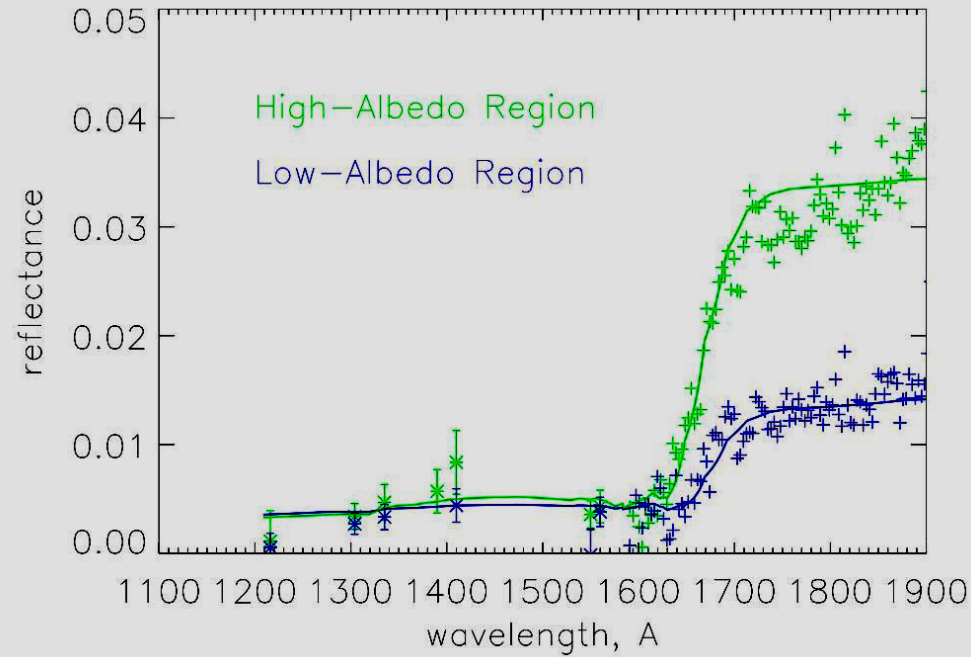
Hyperion & Voyager



Hyperion's Icy Surface

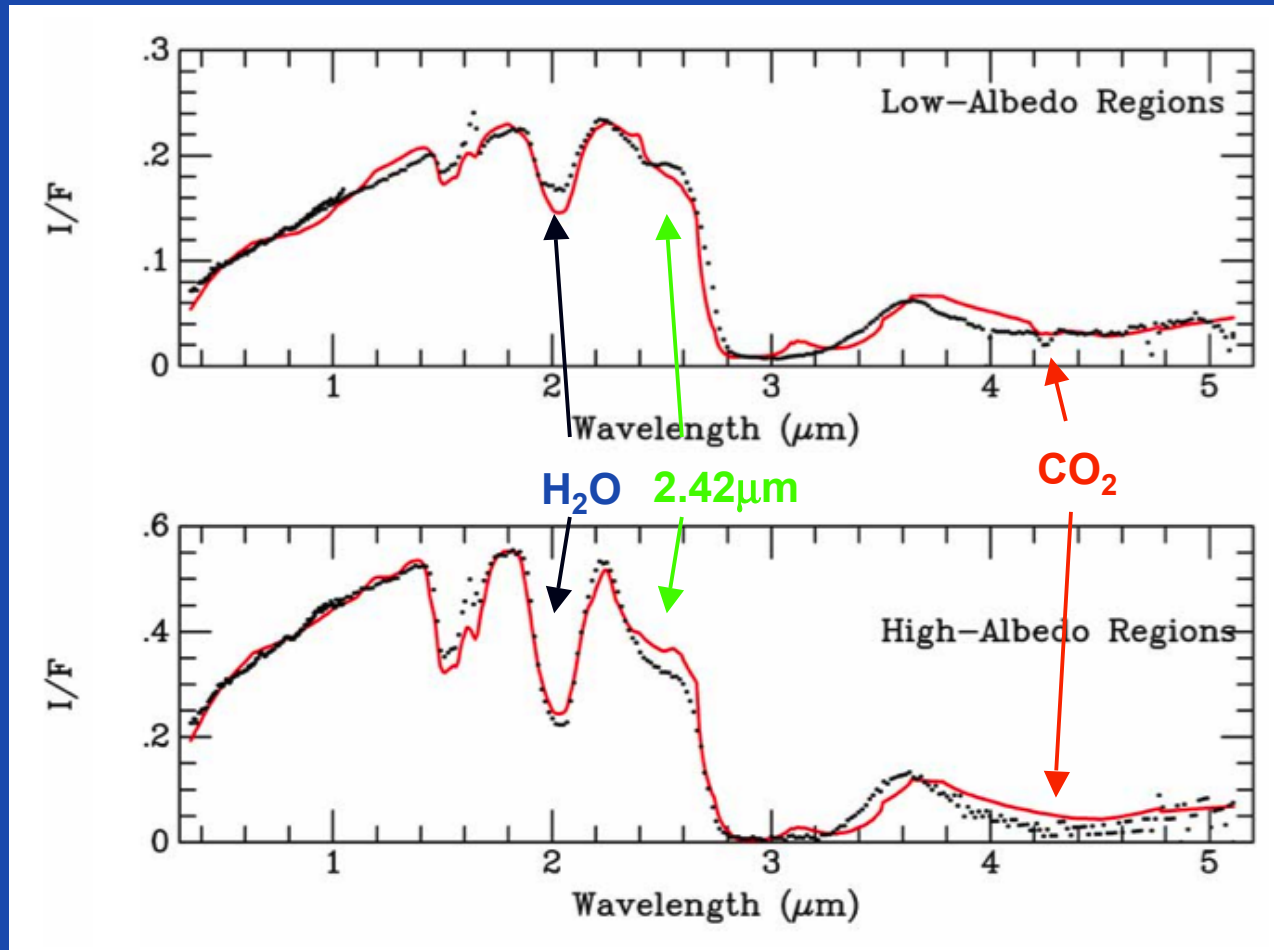


VIMS

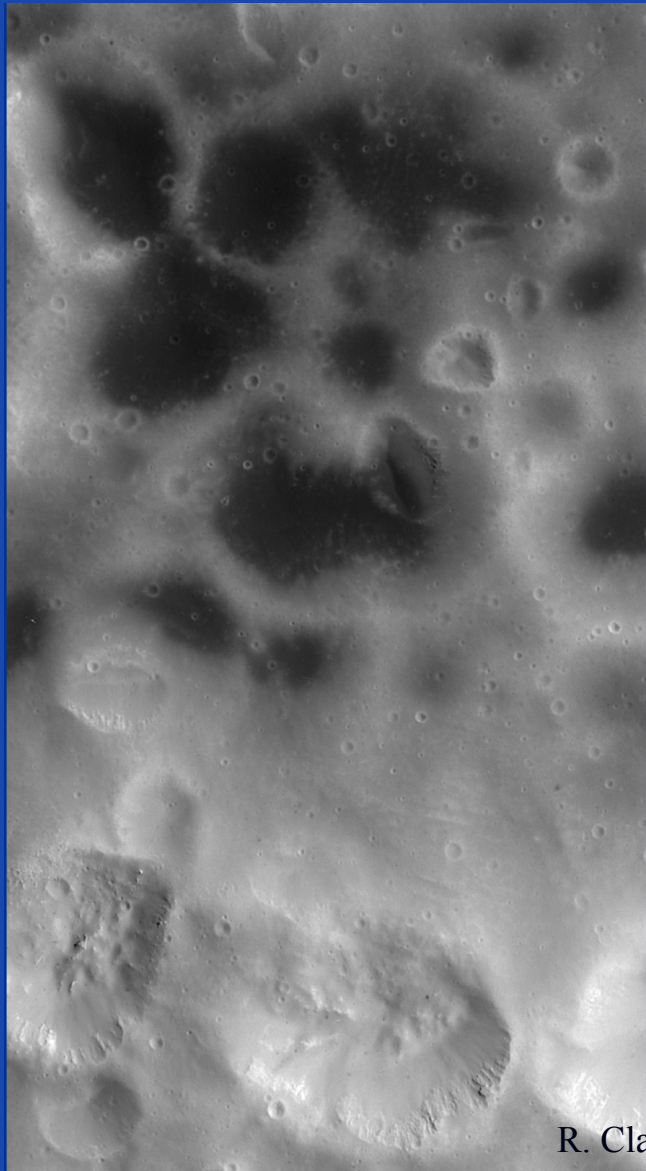


Brightest regions are exposed water ice in the rim of the crater that dominates the hemisphere in view (Hendrix & Hansen, 2007).

Spectra of bright and dark regions

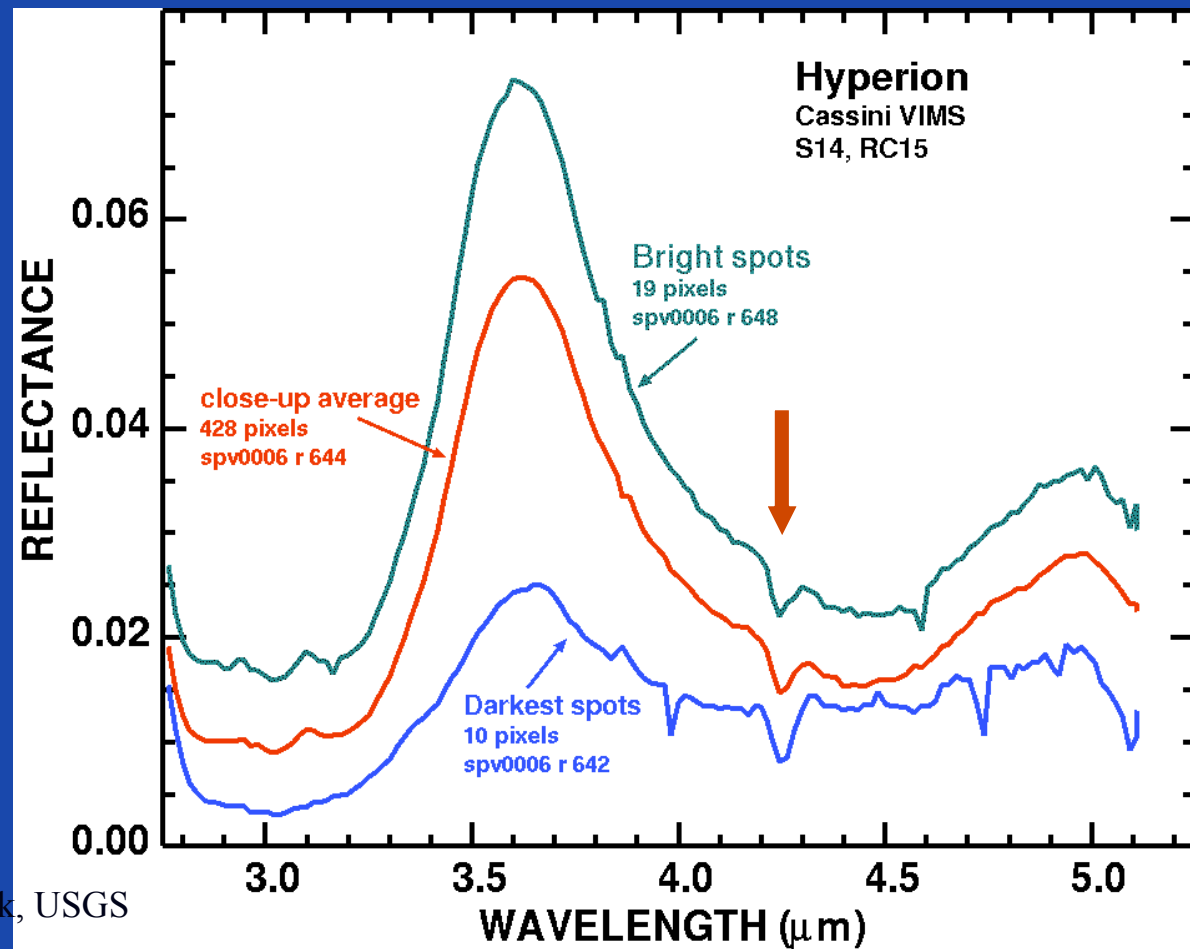
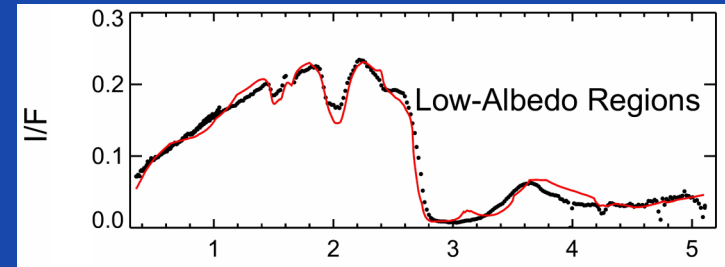


CO₂ on high- and low-albedo regions of Hyperion

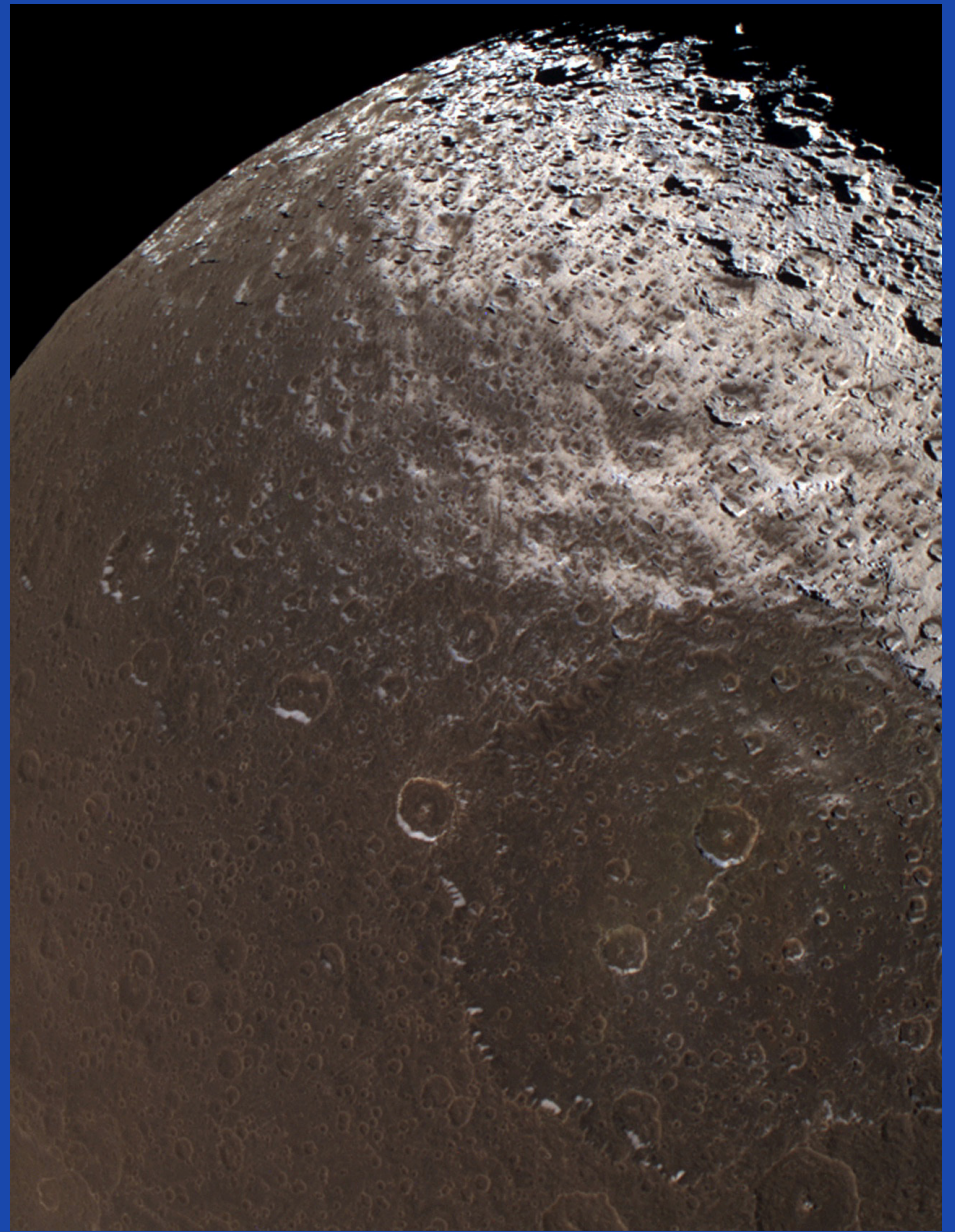
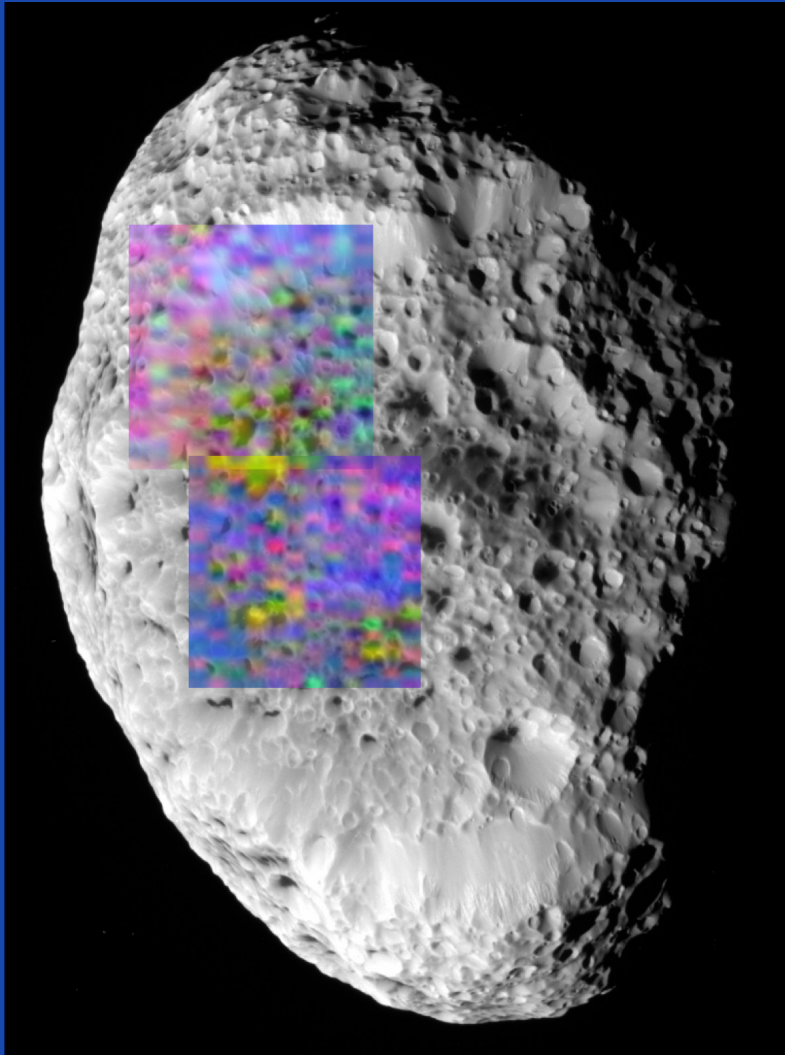


R. Clark, USGS

Full spectrum with model,
Cruikshank et al. 2007

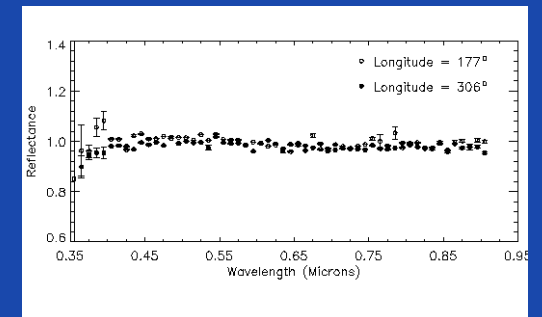
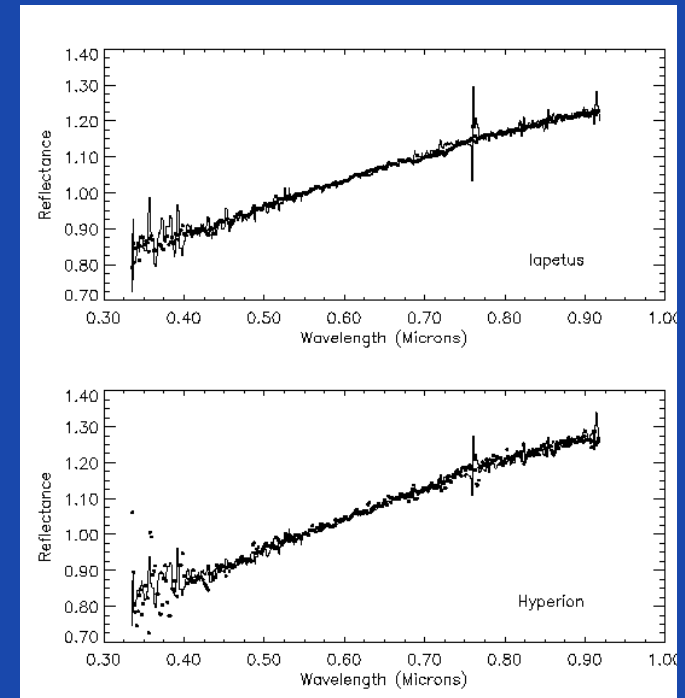
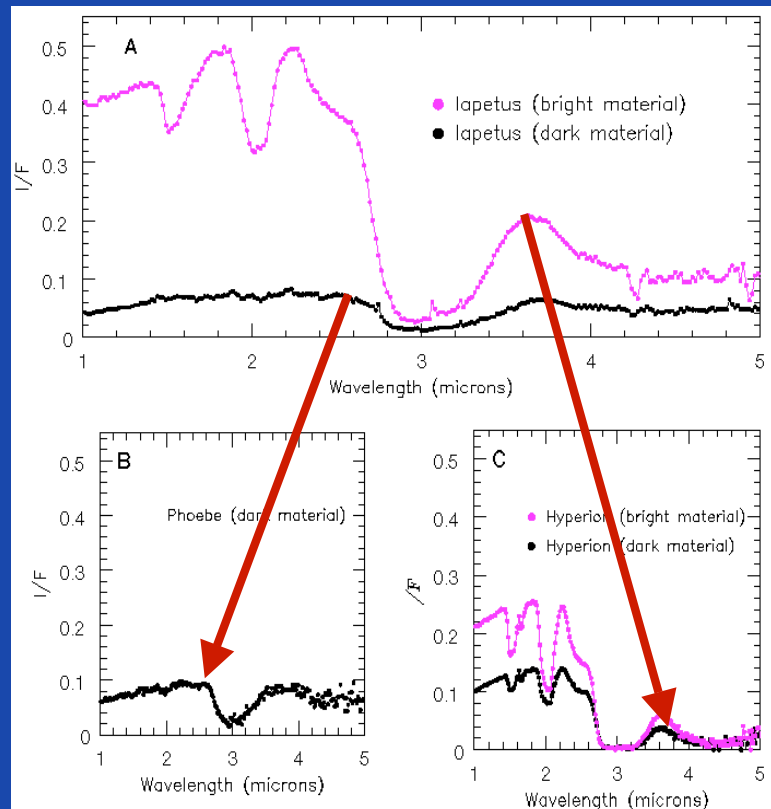


Compositional maps of Hyperion



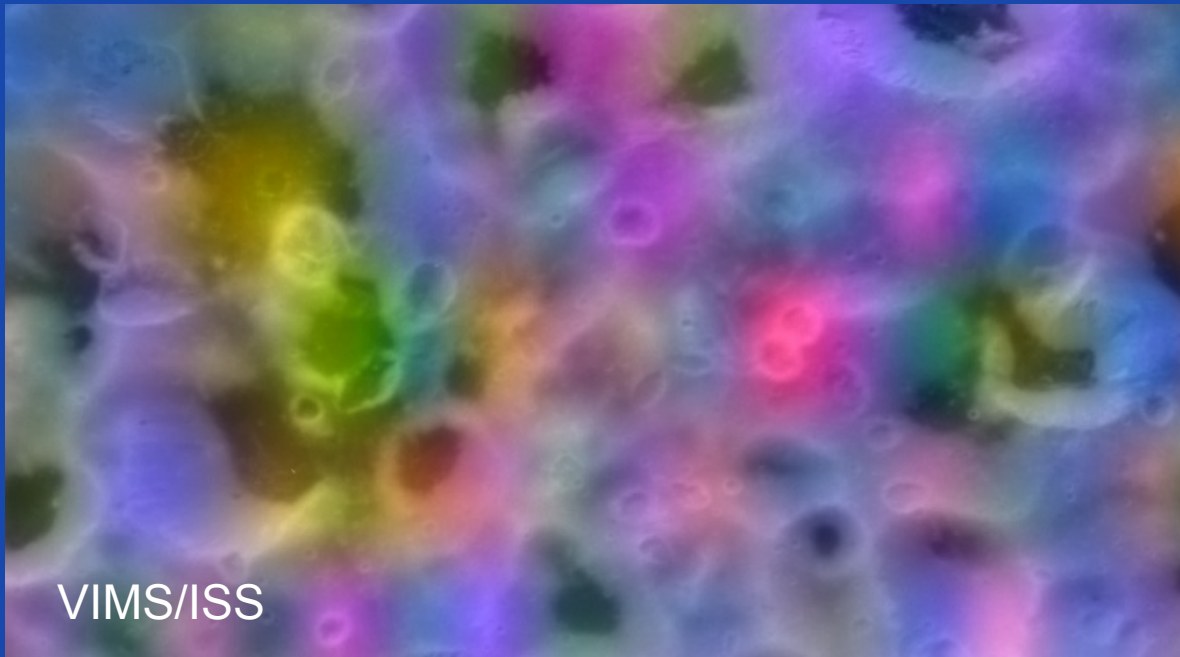
The low-albedo material on Hyperion (in the craters) appears to be the same as that on Saturn's other satellite, Iapetus.

Origin of dark material?



The IR spectra show that Phoebe dark material is similar to Iapetus dark material, but the visual spectra show that Hyperion and Iapetus are more similar.

Buratti et al., 2001, *Icarus* (200-inch Hale telescope)



Hyperion Composition Map

Color code:

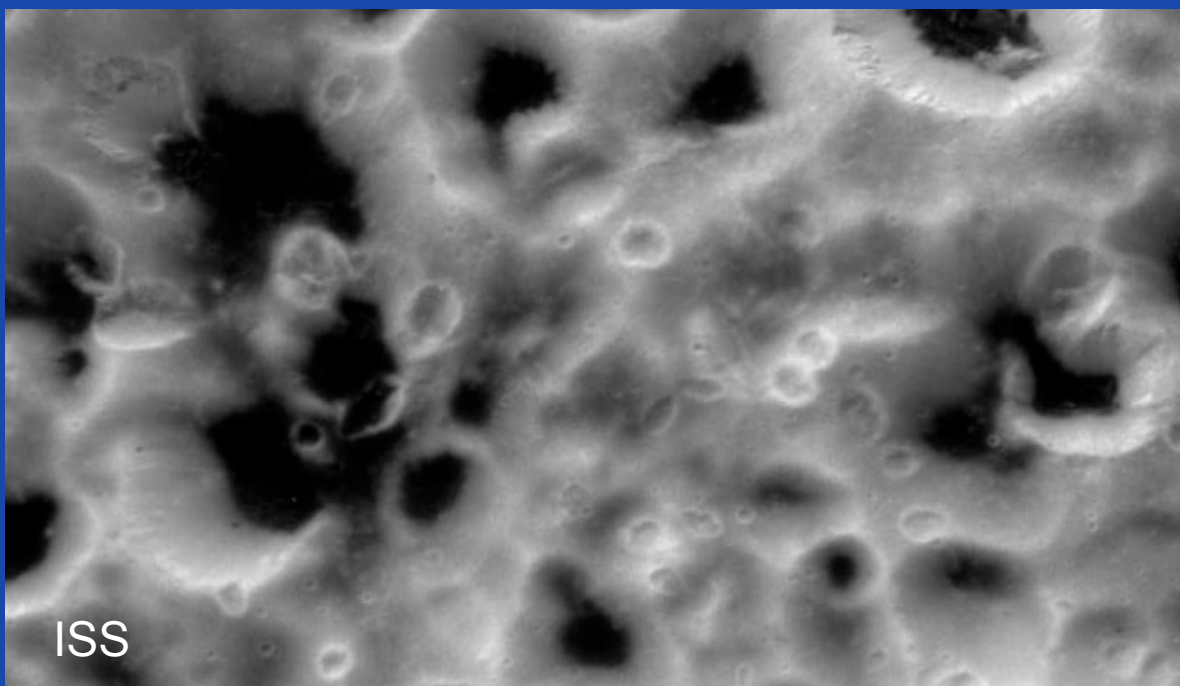
Blue = H₂O band depth

Red = CO₂ band depth

Green = 2.42 μm band

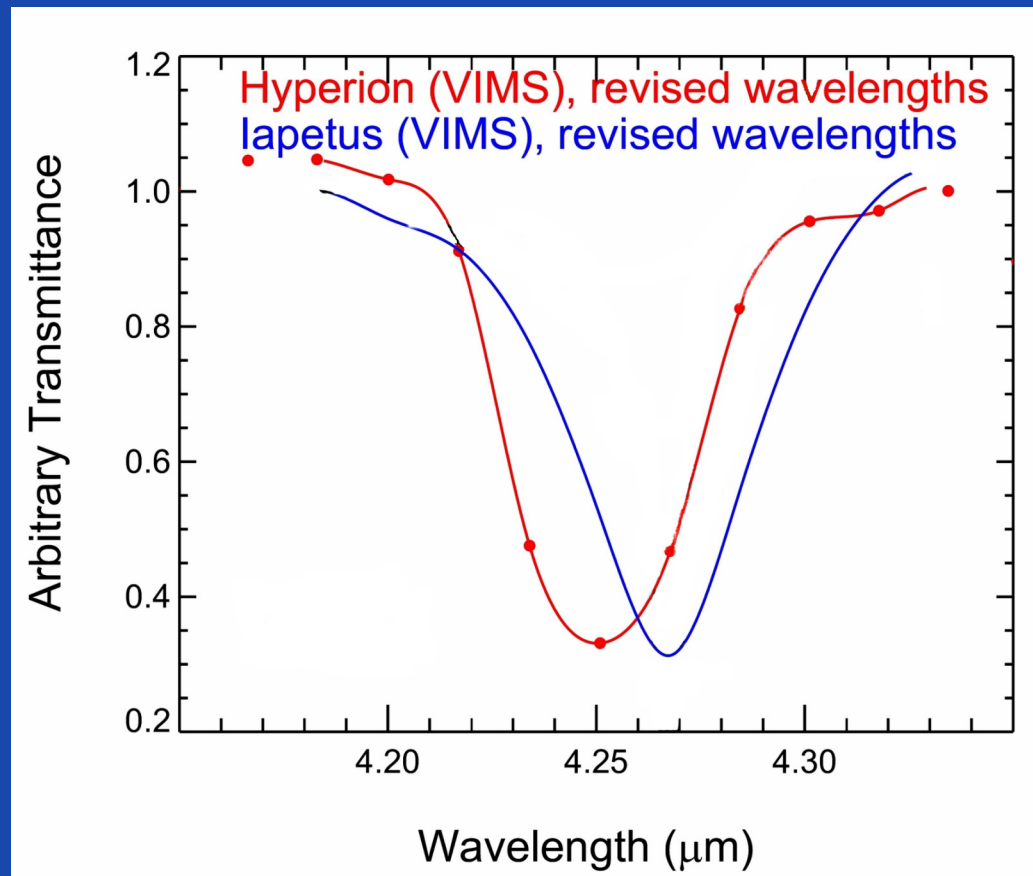
Yellow = CO₂ + 2.42 μm

Magenta = H₂O + CO₂



Map by B. Dalton

CO₂ bands on Iapetus and Hyperion compared.
The shift of the Hyperion's CO₂ band center toward shorter wavelengths indicates that the CO₂ molecules are somehow combined with, or attached to, other materials (possibly water ice molecules).



Origin of the CO₂

■ Original – requires re-supply to surface

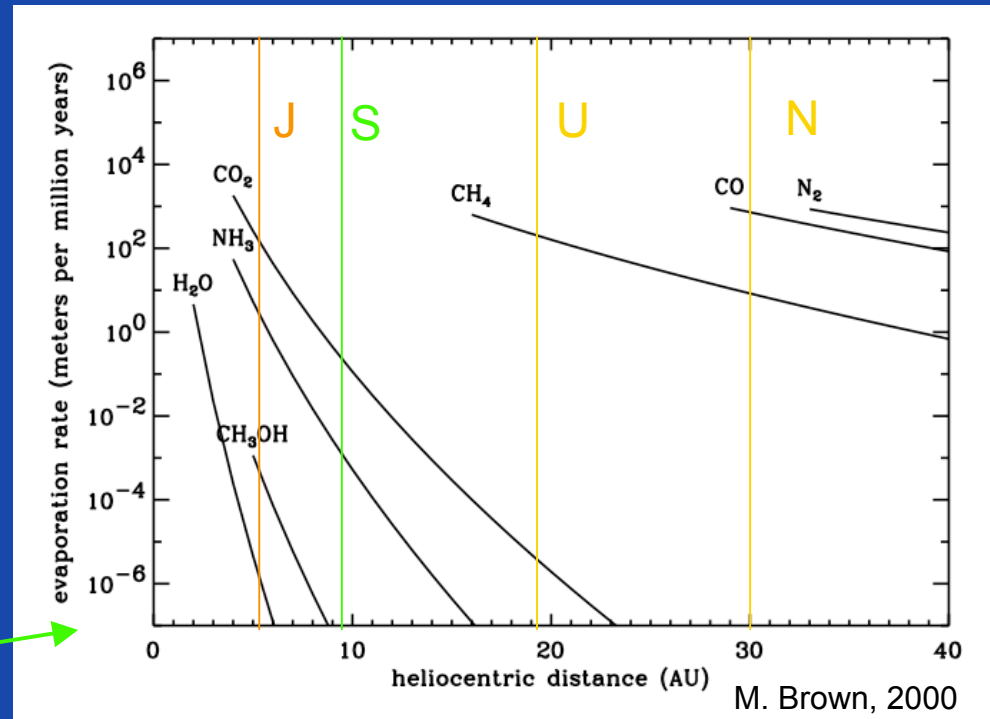
■ Converted CO: $\text{CO} \leftrightarrow \text{CO}_2$ (Moore, Hudson, et al.)

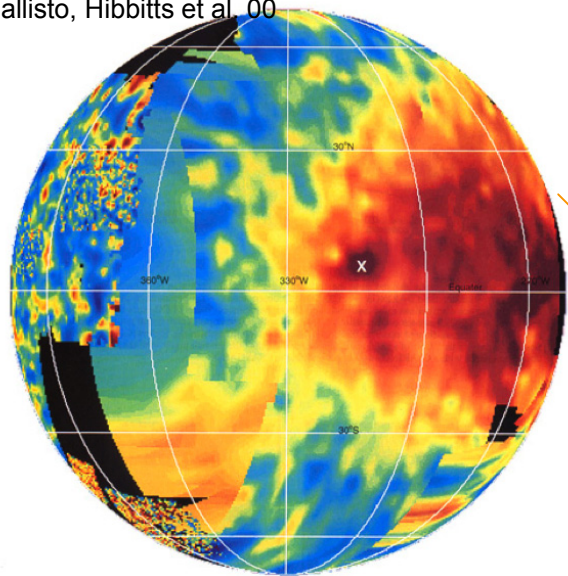
■ Product of H₂O + carbonaceous material with UV

■ Impact or shock-induced chemistry, e.g., $\text{H}_2\text{O} + \text{CH}_3\text{OH} \rightarrow \text{CO}_2$

... (Naa Mvondo et al. 2007)

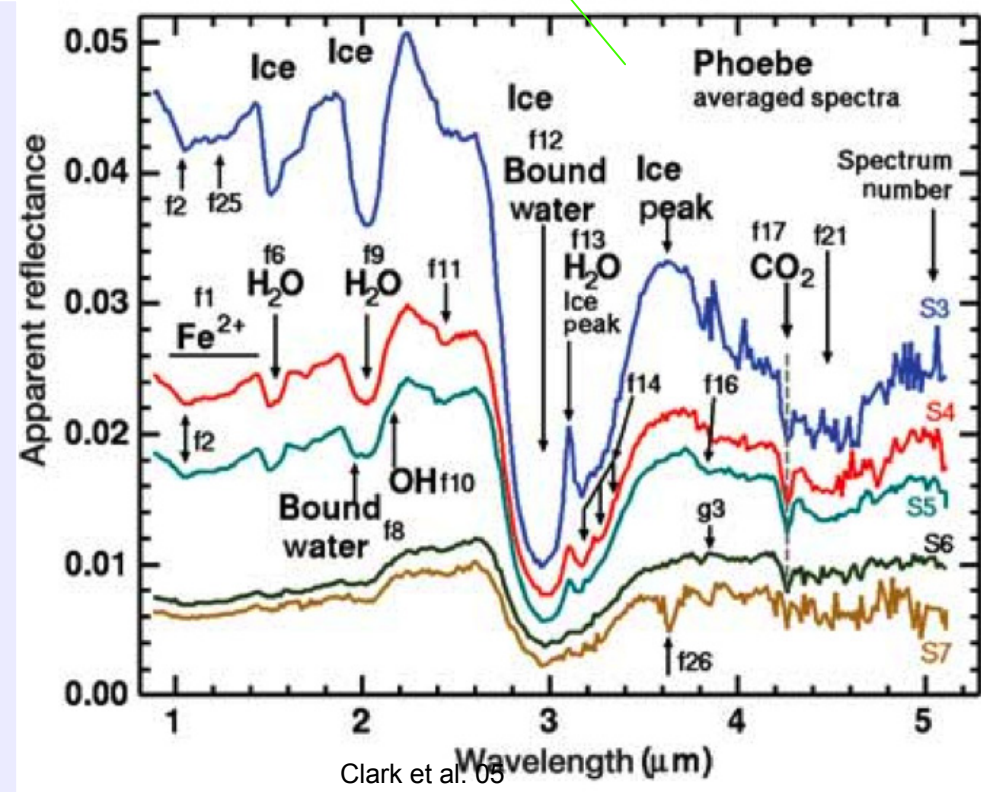
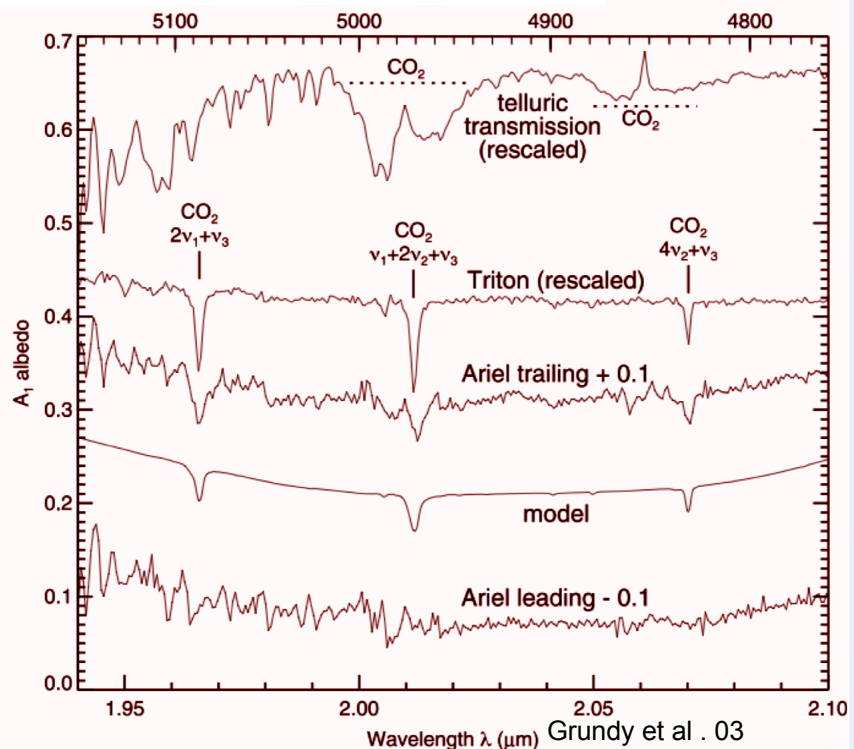
■ Why is Hyperion's CO₂ complexed (when CO₂ on other satellites is not) ?



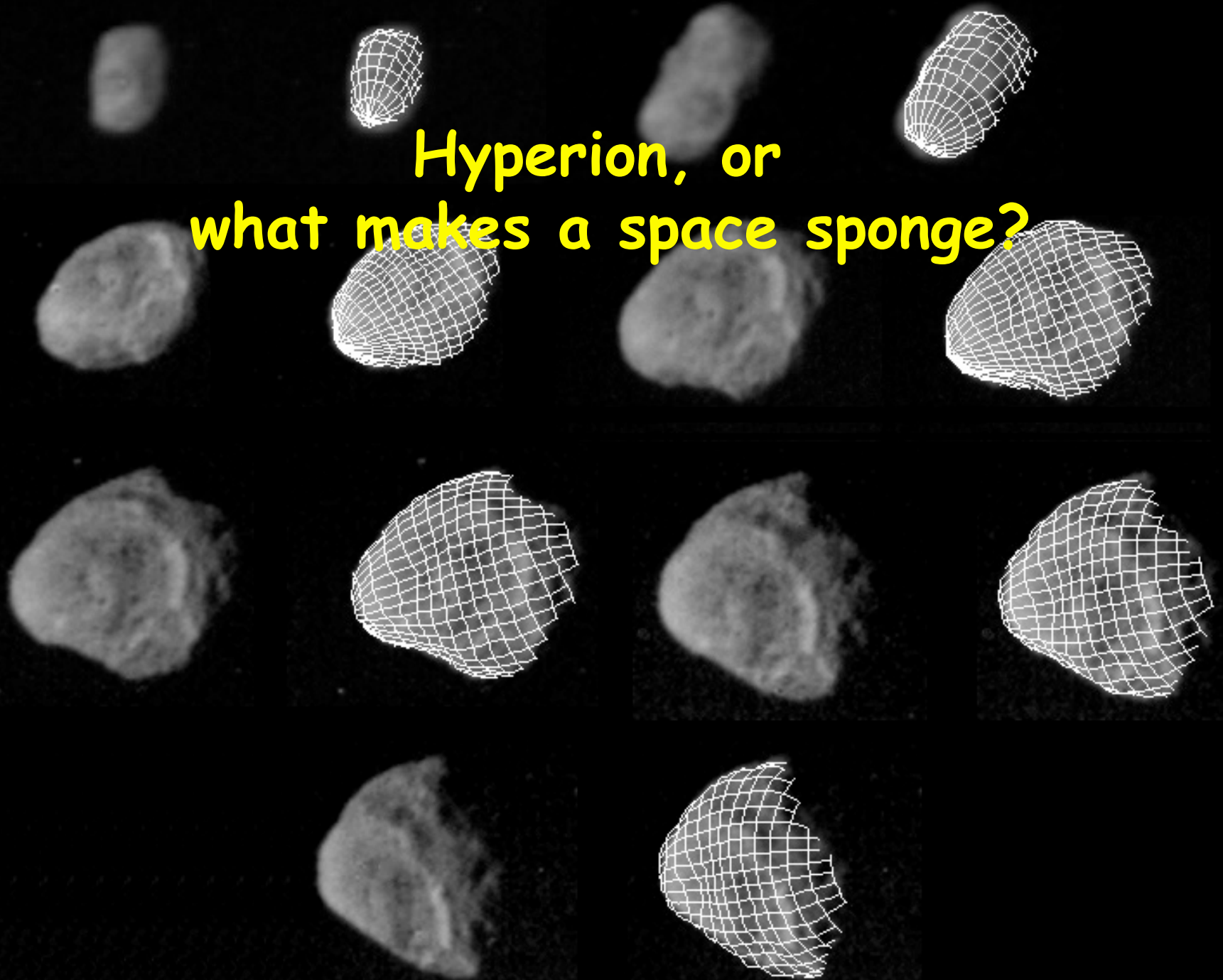


Ubiquitous CO₂?

CO₂ appears to be present on many outer Solar System surfaces including moons of **Jupiter**, **Neptune & Uranus**, and **Saturn**



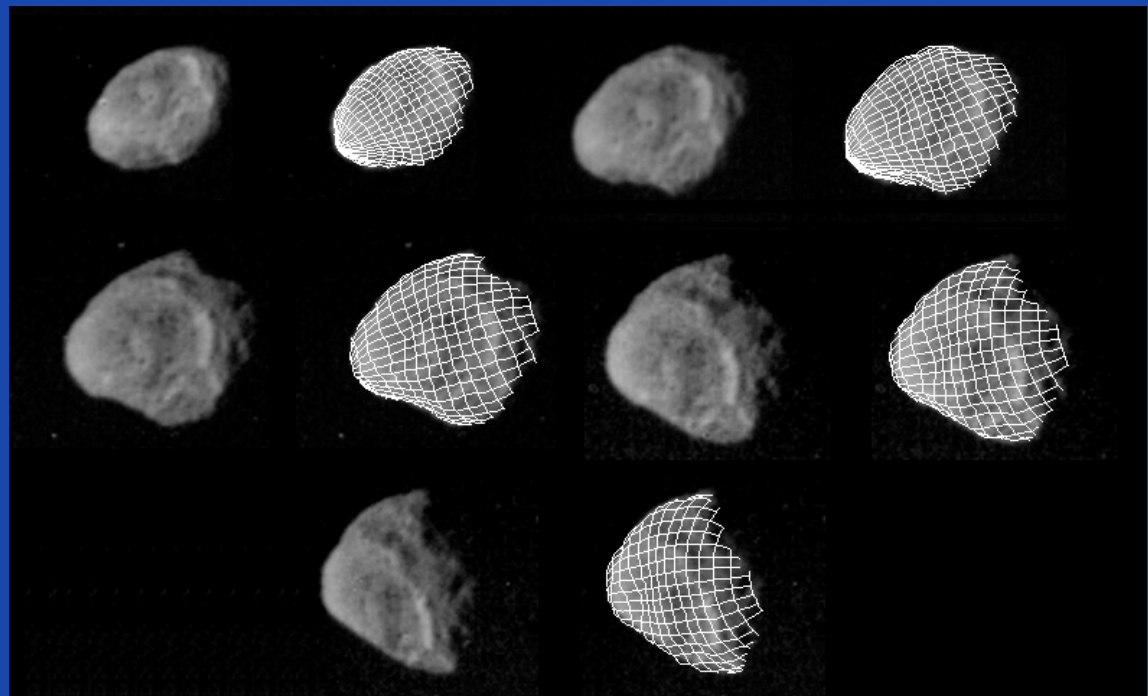
**Hyperion, or
what makes a space sponge?**



Interesting because it has chaotic rotation,
is the largest irregularly-shaped satellite seen to date,
Voyager hinted at spotty albedo distribution.

Cassini made close flyby 26 September 2005, and
Several passes < 500000 km at other times.

Voyager views
of Hyperion



Hyperion: the numbers

Because of chaotic and slow rotation, the generation of the shape model required stitching data from several encounter together.

Mean radius: 135 ± 4 km

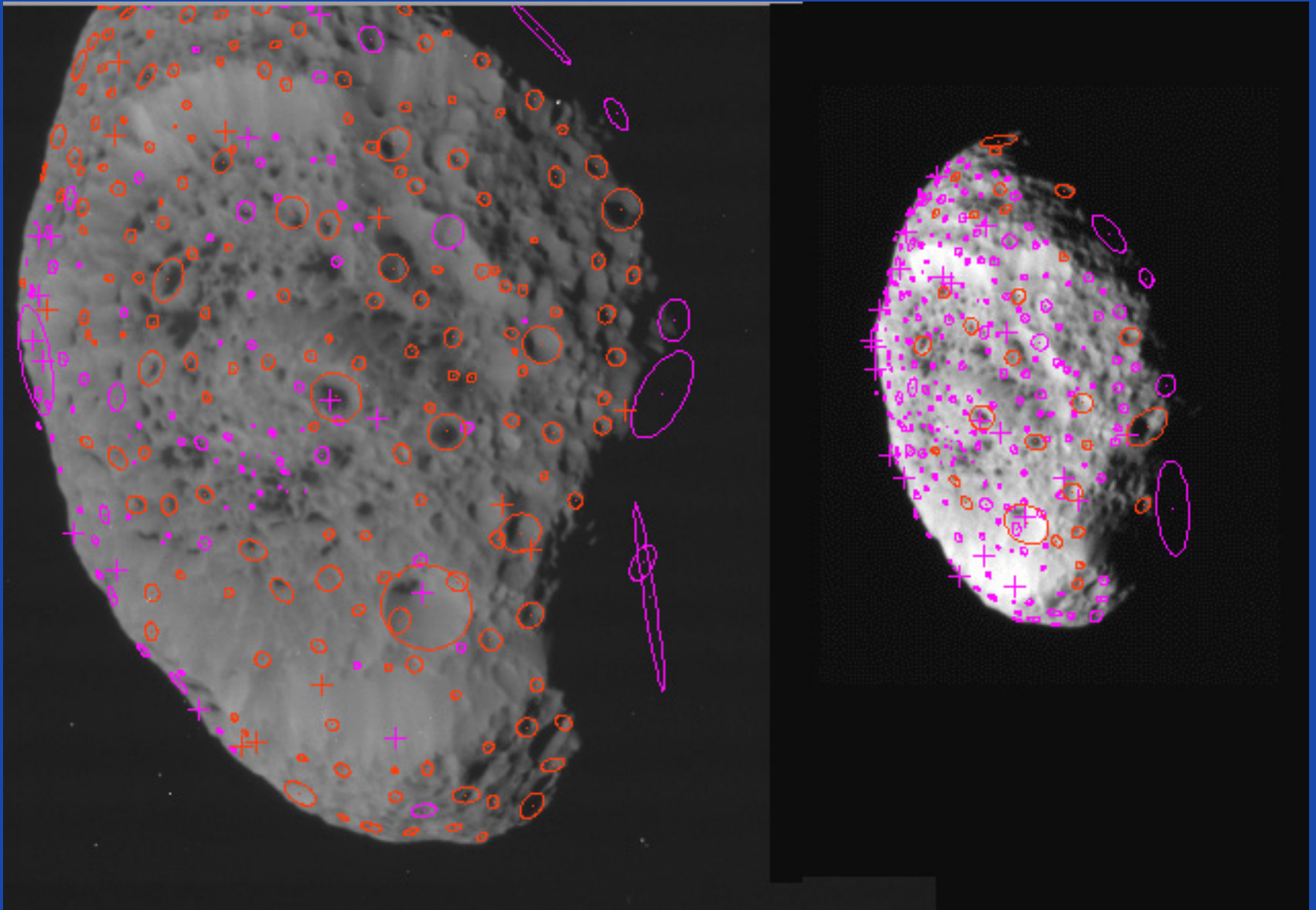
Range of diameters: 195 - 349 km

Mean density 540 ± 50 km m^{-3}

(mass from NAV and RSS teams)

Surface gravity ~ 2 cm s^{-2}

Spin found from observing "control" points



Hyperion: more numbers

UTC range	Instantaneous spin pole solutions:			
	RA	Dec	rate	SS lat
2005,160/18:10:05 - 161/09:58:38	208.4,	35.3	75	-14.1
2005,161/16:43:20 - 162/15:22:01	210.8,	39.3	75	-13.3
2005,228/01:46:59 - 228/20:16:46	271.3,	35.6	72	27.0
2005,228/22:16:13 - 229/12:15:23	266.7,	39.1	72	22.1
2005,268/04:29:06 - 268/17:20:16	315.1,	56.6	72	13.3
2005,268/17:20:16 - 269/02:18:20	299.8,	58.7	72	11.3
2006,179/02:24:03 - 180/04:15:09	184.5,	34.8	72	-43.7

Rate: spin rate in degrees per day

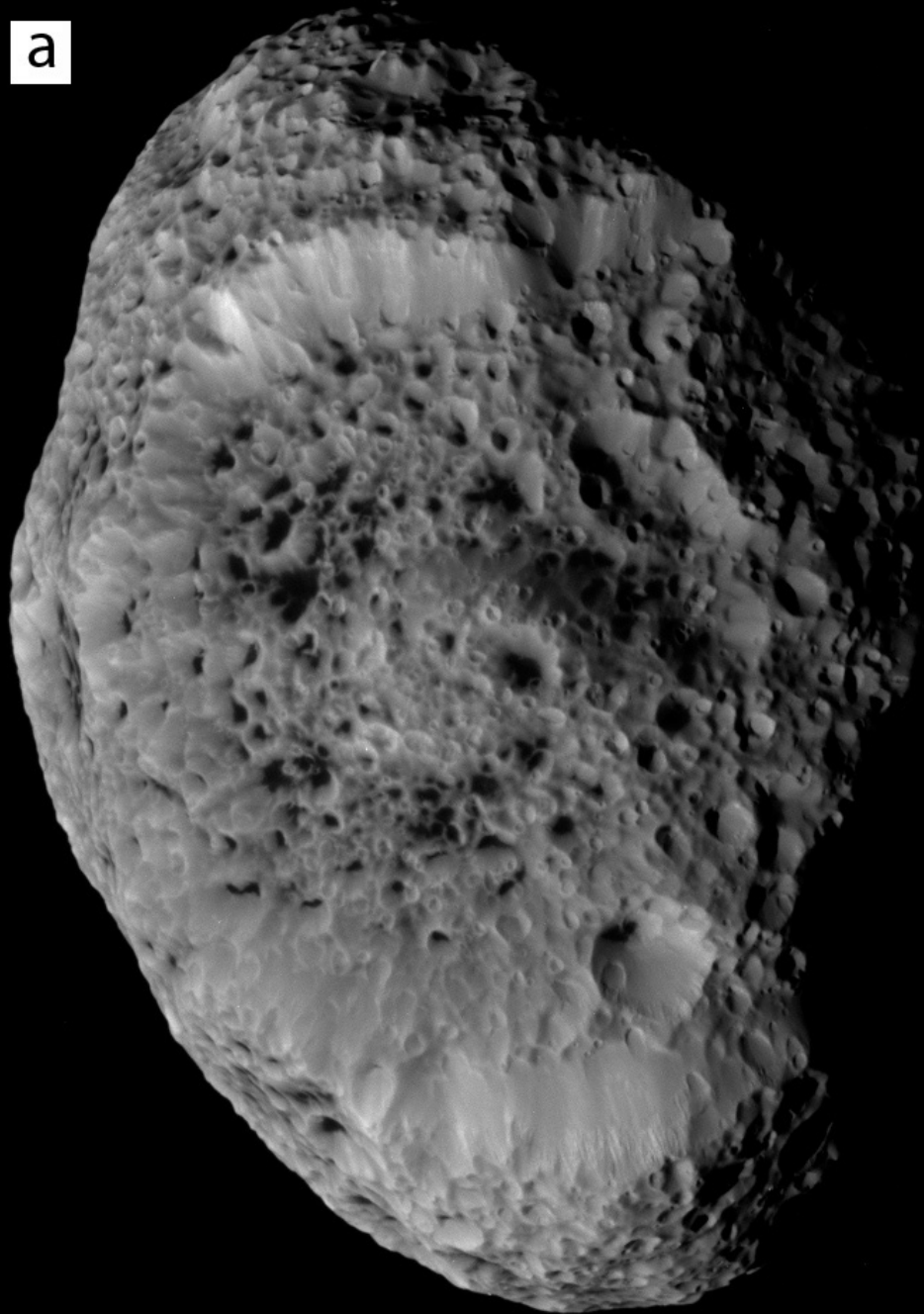
SS lat = sub solar latitude: note drastic changes

Hyperion: the numbers mean what?

Mean density of 540 kgm^{-3} , combined with probability of a lot of water ice, mean porosity is $>40\%$. If there are denser components, porosity is even higher.

Very small ring-related satellites have even lower densities.

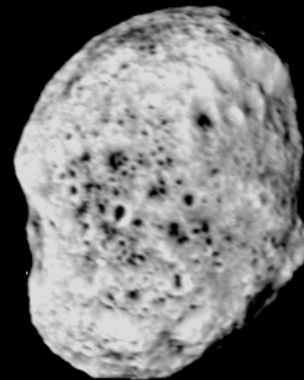
a

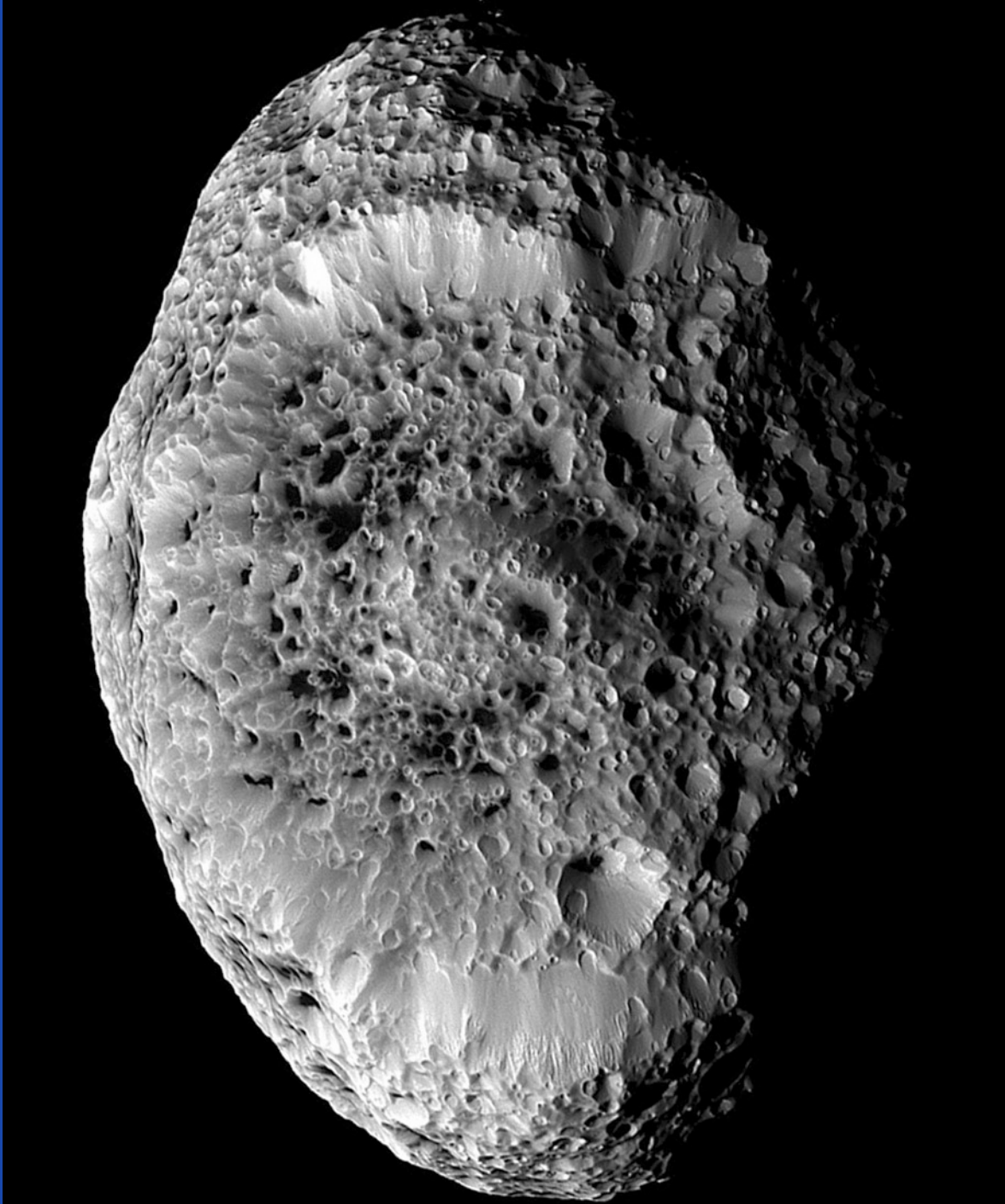


b



c





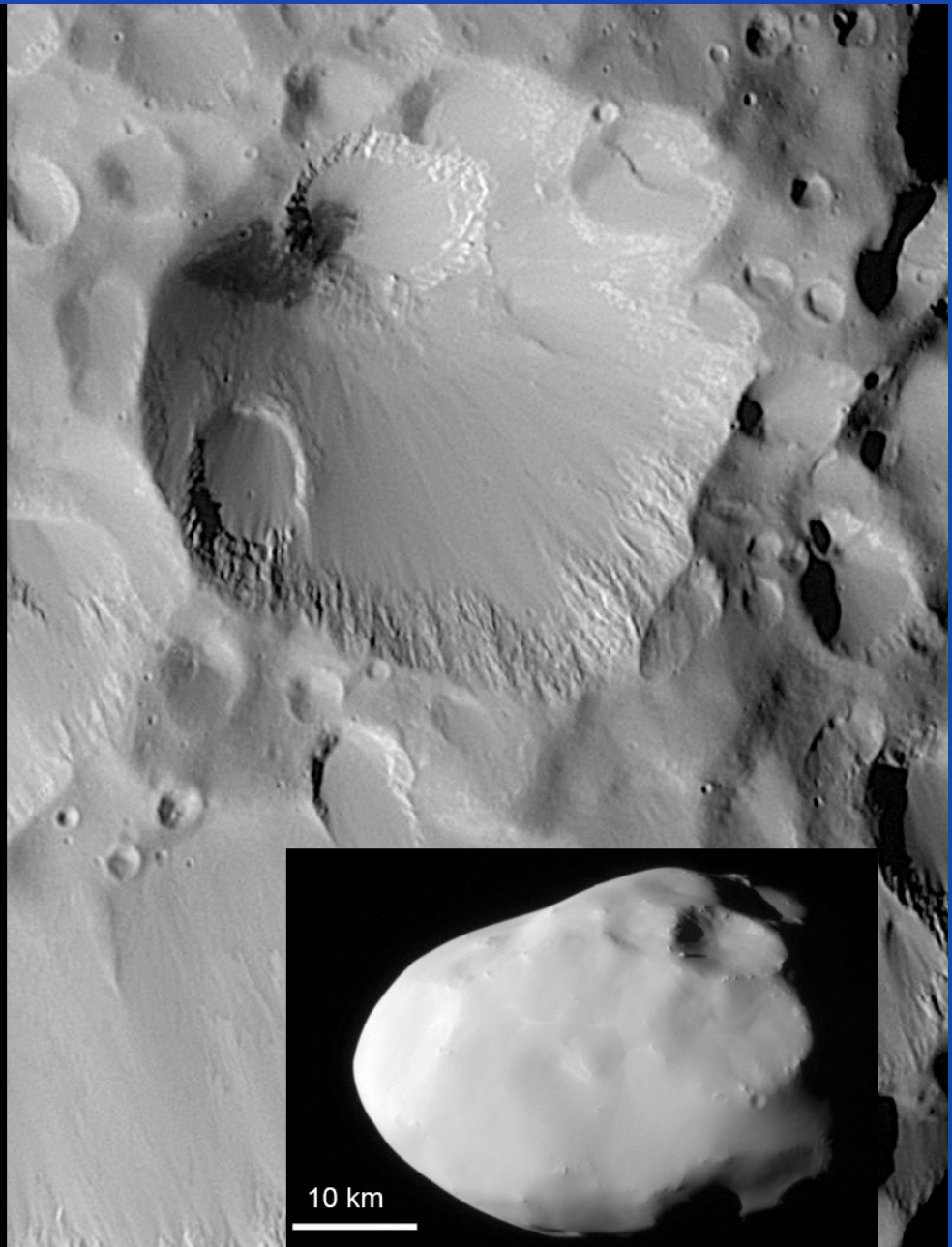
The Sponge

Count obvious
craters; then,
what's left?

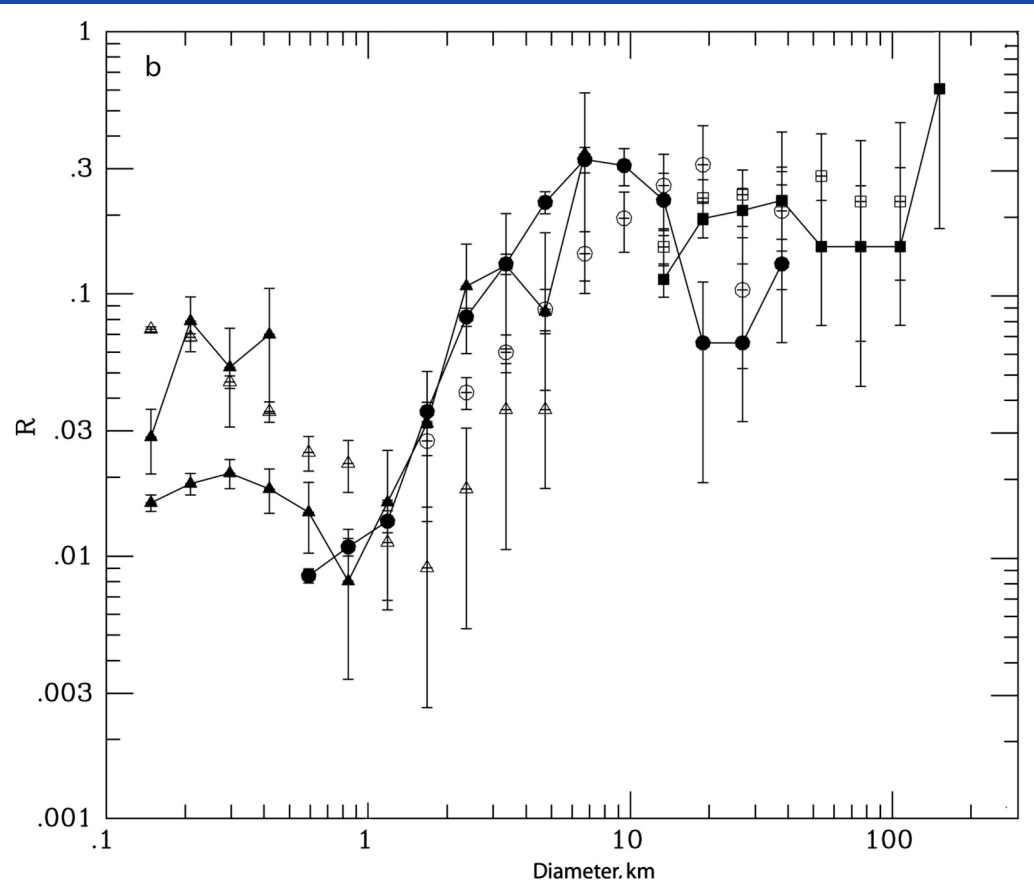
Not much



Phoebe, Hyperion
Telesto at similar
scale



Craters on Hyperion and Phoebe



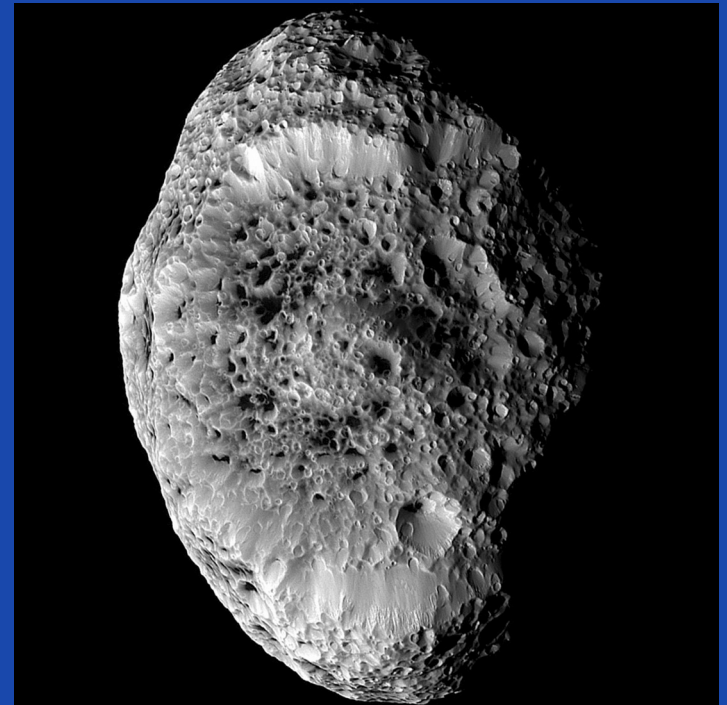
Between 2-10 km,
2x craters on Hyperion
as Phoebe.

>10 km Saturated

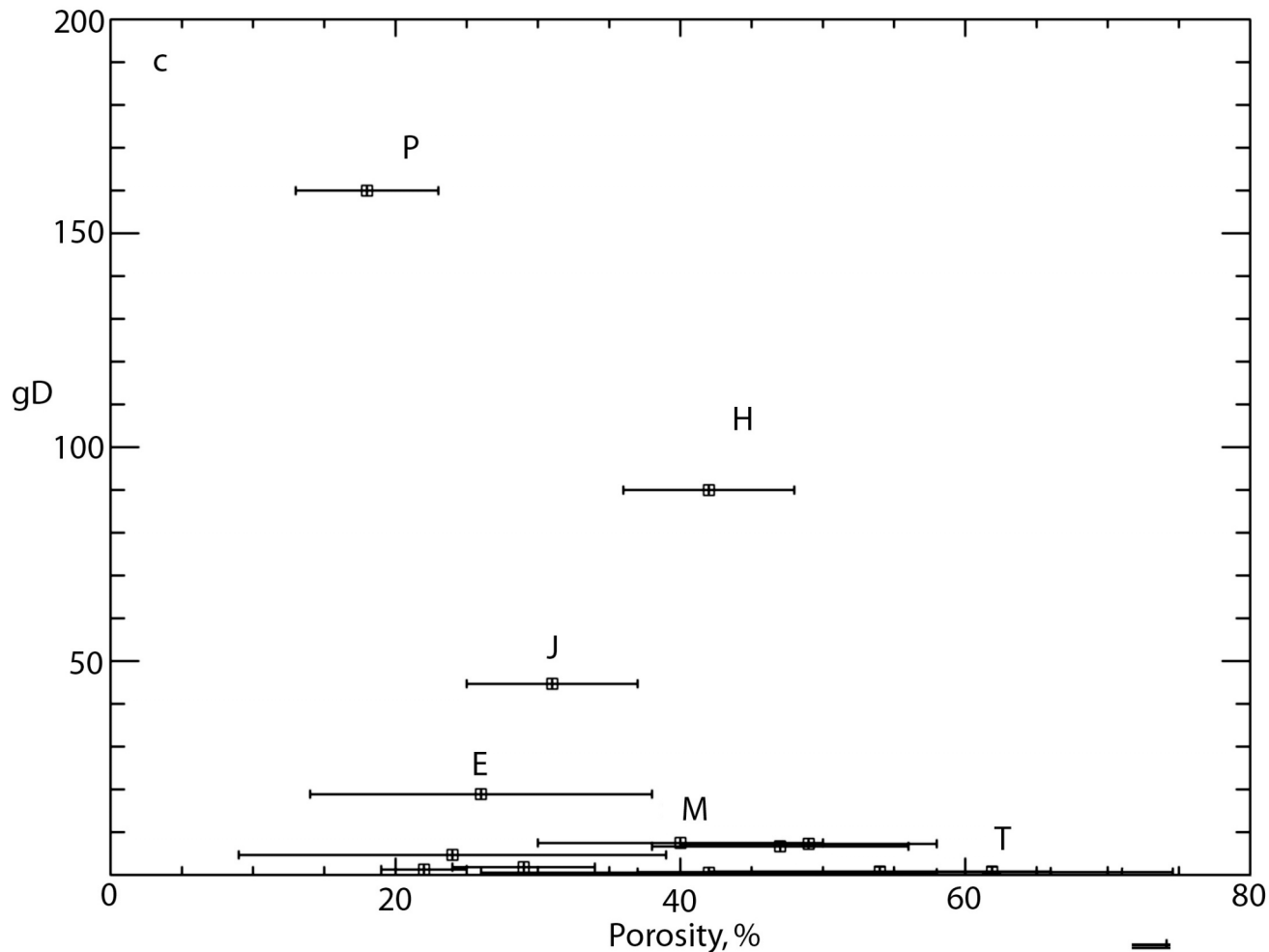
Small: more on Phoebe?



Little fill or erosion in
old craters.
Little sublimation
Little ejecta fill

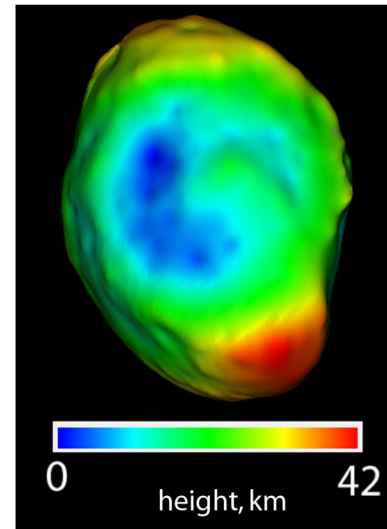
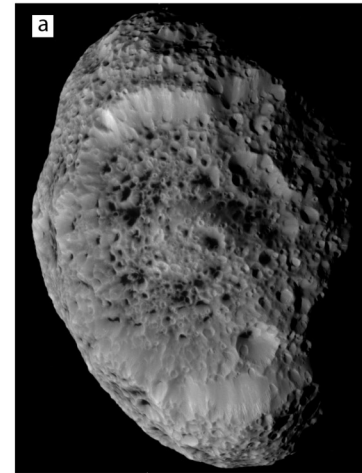
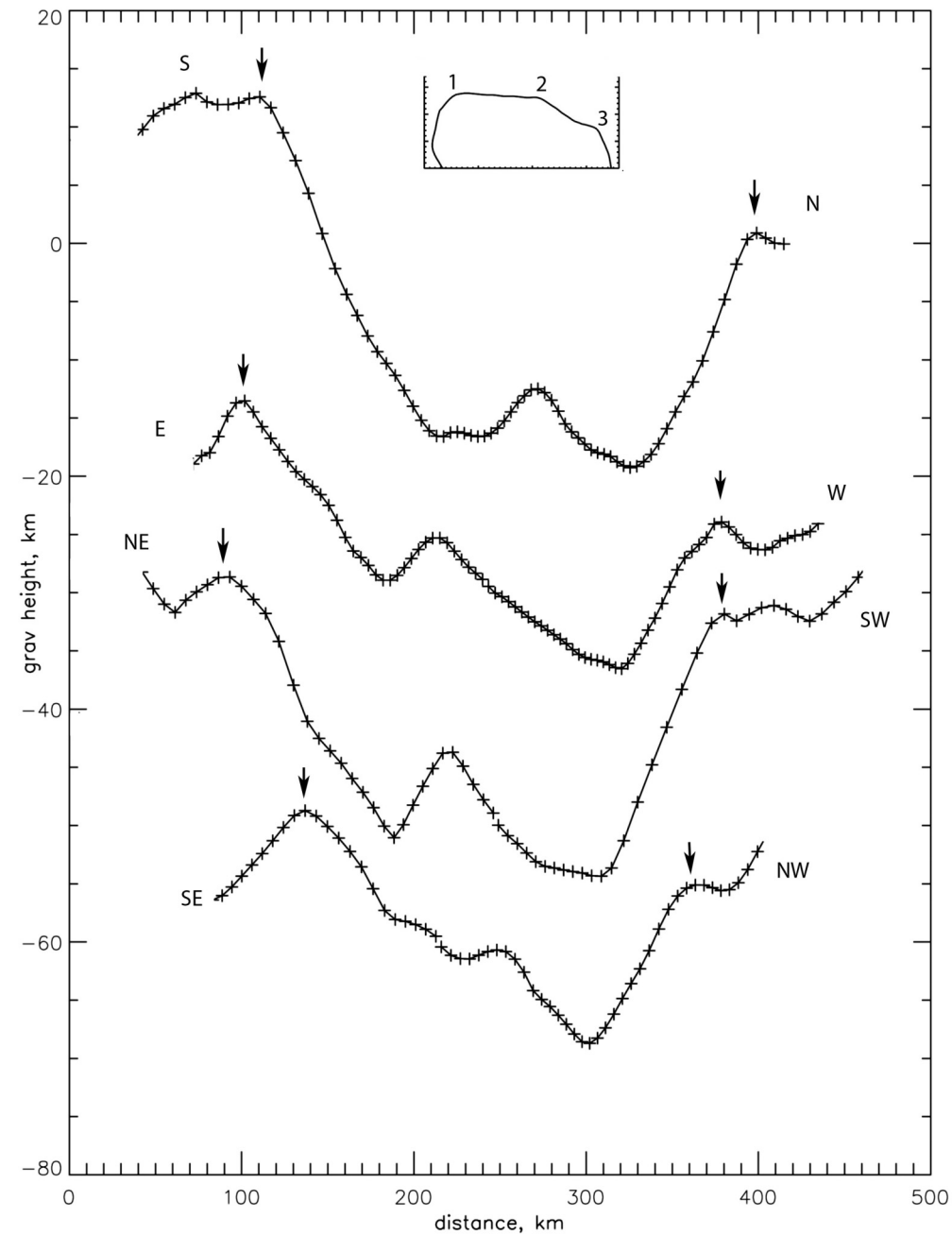


Role of porosity?
Housen and Holsapple, 2003 suggest
Less ejecta for porosities >30%;
and proportional $1/gD$



Very plausible,
if...

Effects of big
impacts?



Geology
Central peaked
crater?
Barely...

Hyperion

Sponge-like appearance from lots of 2-10 km craters with good rim preservation.

No suncup role.

Scant evidence of ejecta, and some evidence for relative lack of ejecta.

Dark material concentrated by downslope motion and/or some sublimation

Porosity effects on craters?